THE Observer's Handbook For 1939

PUBLISHED BY

The Royal Astronomical Society of Canada

C. A. CHANT, EDITOR F. S. HOGG, Assistant Editor david dunlap observatory



THIRTY-FIRST YEAR OF PUBLICATION

TORONTO 198 College Street Printed for the Society By the University of Toronto Press 1939

1939	CALENDAR	1939
JANUARY Sun. 1 8 15 22 29 Mon. 2 9 16 23 30 Tues. 3 10 17 24 31 Wed. 4 11 18 25 Thur. 5 12 19 26 Fri. 6 13 20 27 Sat. 7 14 21 28	FEBRUARY MARCH Sun. 5 12 19 26 Mon. 6 13 20 27 Mon. 6 13 20 27 Mon. 7 14 21 28 Wed. 1 8 15 22 Thur. 2 9 16 3 Fri. 3 10 17 24 Sat. 4 11 18 25 Sat. 4 11 18	19 26 Sun. 2 9 16 23 30 20 27 Mon. 3 10 17 24 21 28 Tues. 4 11 18 25 22 29 Wed. 5 12 19 26 23 30 Thur. 6 13 20 27 24 31 Fri 7 14 21 28
MAY Sun 7 14 21 28 Mon 1 8 15 22 29 Tues. 2 9 16 23 30 Wed. 3 10 17 24 31 Thur. 4 11 18 25 Fri. 5 12 19 26 Sat. 6 13 20 27	$\begin{array}{c c c c c c c c c c c c c c c c c c c $	24 31 Mon. 7 14 21 28 25 Tues. 1 8 15 22 29 26 Wed. 2 9 16 23 30 27 Thur. 3 10 17 24 31 28 Fri. 4 11 18 25
SEPTEMBER Sun. 3 10 17 24 Mon 4 11 18 25 Tues. 5 12 19 26 Wed. 6 13 20 27 Thur. 7 14 21 28 Fri. 1 8 15 22 29 Sat. 2 9 16 23 30	OCTOBER NOVEMB Sun. 1 8 15 22 29 Sun. 5 12 Mon 2 9 16 23 30 Mon. 6 13 Tues. 3 10 17 24 31 Tues. 7 14 Wed. 4 11 18 25 Wed. 18 15 Thur. 5 12 19 26 Thur. 2 9 16 Fri. 6 13 20 27 Fri. 3 10 17 Sat. 7 14 21 28 Sat. 4 11 18	19 26 Sun. 3 10 17 24 31 20 27 Mon. 4 11 18 25 . 21 28 Tues. 5 12 19 26 . 22 29 Wed. 6 13 20 27 . 23 30 Thur. 7 14 21 28 . 24 Fri. 1 85 22 29 .

Julian Day Calendar, 1939

J. D. 2,420,000 plus the following:

Jan. 1	May 1	Sept. 1
Feb. 1	June 1	Oct. 1
Mar. 1	July 19446	Nov. 1
Apr. 1	Aug. 1	Dec. 1

The Julian Day commences at noon. Thus J.D. 2,429,265.0 = Jan. 1.5 G.C.T.

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PREFACE

In the HANDBOOK for 1939, which is the thirty-first issue, numerous changes have been made. By giving the times of sunrise and sunset for every second day, although additional latitudes have been included, there has been a saying of six pages of space. This has allowed the inclusion of times of beginning and ending of twilight, of ephemerides of Saturn's satellites and of the brighter asteroids, the extension of information on Meteors and Occultations, and the insertion of a table of miscellaneous Astronomical Data. The table of Satellites has also been revised. It was hoped to revise the table of The Brightest Stars, but circumstances prevented it this year.

The small star maps at the back necessarily contain only a few objects. Four similar maps 9 inches in diameter are obtainable from the Director of University Extension, University of Toronto, for one cent each. Observers desiring fuller information are recommended to obtain Norton's Star Atlas and Reference Handbook (Gall and Inglis, price 12s 6d; supplied also by Eastern Science Supply Co., Boston). The sixth edition contains late information.

The changes enumerated have been possible through the help of Dr. F. S. Hogg who is now Assistant Editor. Cordial thanks are tendered to all who assisted but especially to Miss Ruth J. Northcott, M.A., of the David Dunlap Observatory.

C. A. CHANT.

David Dunlap Observatory, Richmond Hill, Ont., December 1938.

ANNIVERSARIES AND FESTIVALS 1939

New Year's DaySun. Jan. 1	Corpus ChristiThu. June 8
EpiphanyFri. Jan. 6	St. John Baptist (Midsummer
Septuagesima SundayFeb. 5	Day)
Quinquagesima (Shrove	Dominion DaySat. July 1
Sunday)	Birthday of Queen Elizabeth
Ash WednesdayFeb. 22	(1900)Fri. Aug. 4
Quadragesima (First Sunday	Labour Day Mon. Sept. 4
in Lent)	Hebrew New Year (Rosh
St. David	Hashanah)Thu. Sept. 14
St. PatrickFri. Mar. 17	St. Michael (Michaelmas
Annunciation (Lady	Day)Fri. Sept. 29
Day)Sat. Mar. 25	All Saints' Day Wed. Nov. 1
Palm SundayApr. 2	Remembrance DaySat. Nov. 11
Good FridayApr. 7	St. Andrew
Easter SundayApr. 9	First Sunday in Advent Dec. 3
St. GeorgeSun. Apr. 23	Accession of King George VI
Rogation Sunday May 14	(1936)Mon. Dec. 11
Ascension Day	Birthday of King George VI
Empire Day (Victoria	(1895), Thu. Dec. 14
Day)	Christmas Day Mon. Dec. 25
Birthday of the Queen Mother,	
Mary (1867)Fri. May 26	
Pentecost (Whit Sunday) May 28	Thanksgiving Day, date set by
Trinity SundayJune 4	Proclamation

SYMBOLS AND AGBREVIATIONS

SIGNS OF THE ZODIAC

Υ Aries 0°	ର Leo120°	オ Sagittarius240 ^e
V Taurus 30°	WP Virgo 150°	て Capricornus 270°
		Aquarius 300°
6 Cancer	M Scorpio 210°) (Pisces

SUN, MOON AND PLANETS

\odot The Sun.	• The Moon generally.	24 Jupiter.
New Moon.	g Mercury.	b Saturn.
S Full Moon.	Q Venus.	б or Ӊ Uranus.
First Quarter	\oplus Earth.	Ψ Neptune.
C Last Quarter.	J Mars.	P Pluto

ASPECTS AND ABBREVIATIONS

σ' Conjunction, or having the same Longitude or Right Ascension. ∂ Opposition, or differing 180° in Longitude or Right Ascension. □ Quadrature, or differing 90° in Longitude or Right Ascension. Ω Ascending Node; ♡ Descending Node. a or A.R., Right Ascension; ∂ Declination. h, m, s, Hours, Minutes, Seconds of Time. °'", Degrees, Minutes, Seconds of Arc.

THE GREEK ALPHABET

Α, α,	Alpha.	Ι,ι,	Iota.	Ρ,ρ,	Rho.
Β, β,	Beta.	Κ, κ,	Kappa.	Σ, σ, ς,	Sigma.
Γ,γ,	Gamma.	Λ, λ,	Lambda.	Τ, τ,	Tau.
$\Delta, \delta,$	Delta.	Μ, μ,	Mu.	Υ, ν,	Upsilon,
Ε, ε,	Epsilon.	Ν, ν,	Nu.	Φ, φ,	Pĥi.
Ζ,ζ,	Zeta.	$\Xi, \xi,$	Xi.	Χ, χ,	Chi.
Η, η,	Eta.	0,0,	Omicron.	Ψ,ψ,	Psi.
θ,θ,ϑ,	Theta.	Π,π,	Pi.		Om ega.

THE CONFIGURATIONS OF JUPITER'S SATELLITES

In the Configurations of Jupiter's Satellites (pages 27, 29, etc.), O represents the disc of the planet, d signifies that the satellite is on the disc, * signifies that the satellite is behind the disc or in the shadow. Configurations are for an inverting telescope.

ABBREVIATIONS FOR THE CONSTELLATIONS

Andromeda	. And	Andr
Antlia		Antl
Apus	.Aps	Apus
Aquarius	. Agr	Aqar
Aquila.	Adl	Aqil
Ara	Aro	Arae
Ara		
Aries	. Ari	Arie
Auriga	. Aur	Auri
Bootes		Boot
Coolum	Coo	Cael
	. Cae	
Caelum	. Cam	Caml
Cancer.	. Cnc	Canc
Canes Venatici	CVn	CVen
Canie Maion	CMa	CMai
Canis Major Canis Minor	. Civia	CMaj
Canis Minor	.CM1	CMin
Capricornus	. Can	Capr
Carina		Cari
Cassiopeia		Cass
Centaurus	. Cen	Cent
	. Cep	Ceph
Cotus	. Cet	Ceti
	. Cel	
	. Cha	Cham
Circinus	. Cir	Circ
Columba		Colm
Como Denorioso	Con	
Coma Berenices	. Com	Coma
Corona Australis		CorA
Corona Borealis	. CrB	CorB
Corvus		Corv
Crater		Crat
Crux		Cruc
Cygnus	. Cvg	Cygn
Delphinus	ົ້ມ	Dlph
Dorado		Dora
Draco	. Dra	Drac
Equuleus	. Eau	Equl
Eridanus	Fri	Erid
Engange	. En	
	.For	Forn
Gemini	. Gem	Gemi
Grus	.Gru	Grus
	.Her	Herc
ITercules		
Horologium	. Hor	Horo
Hydra	.Hya	Hyda
Hydrus	Hvi	Hydi
Indue	Ind	Indi
Indus.		
Lacerta		Lacr
Leo	. Leo	Leon
Leo Minor	LMi	LMin
Lepus	. сер	Leps

Libra. Lupus. Lynx. Lynx. Lyra. Mensa. Microscopium. Monoceros. Musca. Norma. Octans. Ophiuchus. Ophiuchus. Ophiuchus. Ophiuchus. Ophiuchus. Ophiuchus. Ophiuchus. Perseus. Perseus. Phoenix. Pisces. Pisces. Pisces. Pisces. Puppis. Pyxis. Reticulum. Sagitta. Sagitta. Sagitta. Scorpius. Sculptor. Scutum. Sextans. Taurus. Telescopium. Sextans. Telescopium. Sextans. Sextans. Sextans. Sagita. Sextans. Sextans. Taurus. Telescopium. Sextans.	Lup Lyn Lyn Mon Mon Mor Oct Per Per Per Pic Sgr Sccl Scr Scr Scr	Libr Lupi Lync Lyra Mens Micr Mono Otho Pavo Pegs Phot Pisc Pisc Sgtr Scul Scut Serp Seau
Sagittarius	. Sgr	
Scorpius	. Scr	
Sculptor	. Scl	
Sevtans	Sev	
Taurus	Tau	Taur
Telescopium	. Tel	Tele
Triangulum	. Tri	Tria
Triangulum Australe	.TrA	TrAu
Tucana Ursa Major	. Tuc	Tucn
Ursa Major	.∪Ma	UMaj
Ursa Minor	. U MI	UMin Velr
Virgo		Verr
Volans		Voln
Vulpecula		Vulp
		· ··· ·

The 4-letter abbreviations are intended to be used in cases where a maximum saving of space is not necessarv.

Sary. From Transactions of the I.A.U., Vol. IV., 1932, page 221. UNITS OF LENGTH $= 10^{-8}$ cm. 1 Angstrom unit 1 micron $= 10^{-4}$ cm. 1 meter $= 10^{2}$ cm. = 3.28084 feet 1 kilometer $= 10^{5}$ cm. = 0.62137 miles 1 mile $= 1.60935 \times 10^5$ cm. = 1.60935 km. 1 astronomical unit = 1.49504 × 1013 cm. = 92,897,416 miles 1 light year $= 9.463 \times 10^{17}$ cm. $= 5.880 \times 10^{12}$ miles = 0.3069 parsecs 1 parsec $= 30.84 \times 10^{17}$ cm. $= 19.16 \times 10^{12}$ miles = 3.259 l.v. $= 30.84 \times 10^{23}$ cm. $= 19.16 \times 10^{18}$ miles $= 3.259 \times 10^{6}$ l.y. 1 megaparsec UNITS OF TIME = 23h 56m 04.09s of mean solar time Sidereal day Mean solar day = $24h \ 03m \ 56.56s$ of sidereal time Synodical month = $29d \ 12h \ 44m$; sidereal month = $27d \ 07h \ 43m$ Tropical year (ordinary) $= 365d \ 05h \ 48m \ 46s$ Sidereal year $=365d \ 06h \ 09m \ 10s$ Eclipse year $=346d \ 14h \ 53m$ THE EARTH Equatorial radius, a = 3963.35 miles; flattening, c = (a-b)/a = 1/297.0Polar radius, b = 3950.01 miles 1° of latitude = $69.057 - 0.349 \cos 2\phi$ miles (at latitude ϕ) 1° of longitude = 69.232 cos ϕ -0.0584 cos 3 ϕ miles Mass of earth = 6.6×10^{21} tons; velocity of escape from $\bigoplus = 6.94$ miles/sec. EARTH'S ORBITAL MOTION Solar parallax = 8.''80; constant of aberration = 20.''47Annual general precession = 50.''26; obliquity of ecliptic = $23^{\circ} 26' 50''$ (1939) Orbital velocity = 18.5 miles/sec.; parabolic velocity at \bigoplus = 26.2 miles/sec. SOLAR MOTION Solar apex, R.A. $18h \ 04m$; Dec. $+ 31^{\circ}$ Solar velocity = 12.2 miles/sec. THE GALACTIC SYSTEM North pole of galactic plane R.A. $12h \ 40m$, Dec. $+ 28^{\circ}$ (1900) Center, 325° galactic longitude, = R.A. 17h 24m, Dec. -30° Distance to center = 10,000 parsecs; diameter = 30,000 parsecs. Rotational velocity (at sun) = 262 km./sec. Rotational period (at sun) = 2.2×10^8 years Mass = 2×10^{11} solar masses EXTRAGALACTIC NEBULAE Red shift =+530 km./sec./megaparsec,=+101 miles /sec./million l.y. RADIATION CONSTANTS Velocity of light = 299.774 km./sec. = 186.271 miles/sec. Solar constant = 1.93 gram calories/square cm./minute Light ratio for one magnitude = 2.512; log ratio = 0.4000Radiation from a star of zero apparent magnitude = 3×10^{-6} meter candles Total energy emitted by a star of zero absolute magnitude = 5×10^{25} horsepower MISCELLANEOUS Constant of gravitation, $G = 6.670 \times 10^{-8}$ c.g.s. units Mass of the electron $m = 9.035 \times 10^{-28}$ gm.; mass of the proton $= 1.662 \times 10^{-24}$ gm. Planck's constant, $h = 6.55 \times 10^{-27}$ erg. sec. Loschmidt's number = 2.705×10^{19} molecules/cu. cm. of gas at N.T.P. Absolute temperature = T° K = T° C +273° = 5/9 (T° F +459°) $\pi = 3.141,592,653,6$ $1 \text{ radian} = 57^{\circ}.2958$ = 3437'.75No. of square degrees in the sky = 206.265''=41.253

1939 EPHEMERIS OF THE SUN AT 0h. GREENWICH CIVIL TIME

	parent R.A.	Corr. to Sundial	Apparent Dec.	Date	Apparent R.A.	Corr. to Sundial	Apparent Dec.
$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{c} 55 & 200 \\ 08 & 300 \\ 08 & 307 \\ 34 & 395 \\ 477 & 355 \\ 477 & 356 \\ 477 & 356 \\ 477 & 356 \\ 477 & 356 \\ 477 & 386 \\ 175 & 226 \\ 175 & 226 \\ 175 & 227 \\ 550 & 400 \\ 125 & 227 \\ 150 & 227 \\ 257 & 388 \\ 100 & 22 \\ 277 & 388 \\ 100 & 22 \\ 277 & 388 \\ 100 & 22 \\ 277 & 388 \\ 100 & 22 \\ 277 & 388 \\ 100 & 22 \\ 277 & 388 \\ 100 & 22 \\ 277 & 388 \\ 100 & 22 \\ 277 & 388 \\ 100 & 22 \\ 277 & 388 \\ 100 & 22 \\ 277 & 100 \\ 271 & 290 \\ 100 & 52 \\ 270 & 200 \\ 175 & 200 \\ 1$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{c} \circ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ $	$ \begin{array}{c} \ \ \ \ \ \ \ \ \ \ \ \ \ $	$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$		$\begin{array}{c} & & & & & & & \\ & & & & & & & \\ & & & & & & & \\ & & & & & & & \\ & & & & & & & \\ & & & & & & & \\ & & & & & & & \\ & & & & & & & \\ & & & & & & & \\ & & & & & & & \\ & & & & & & & \\ & & & & & & & \\ & & & & & & & \\ & & & & & & & \\ & & & & & & & \\ & & & & & \\ & & & & & & \\ & & & & & \\ & & & & & & \\ & & & & & & \\ & & & & & & & \\ & & & & & & & \\ & & & & & & & \\ & & & & & & & \\ & & & & & & & \\ & & & & & & & \\ & & & & & & & \\ & & & & & & & \\ & & & & & & & \\ & & & & & & & \\ & & & & & & & \\ & & & & & & & \\ & & & & & & & \\ & & & & & & & \\ & & & & & & & \\ & & & & & & & \\ & & & & & & & & \\ & &$

Corr. to Sundial gives that quantity which, if the sign be + must be added to, or if the sign be - subtracted from, the *apparent* or sundial time, to get the local *mean* time.

SOLAR AND SIDEREAL TIME

In practical astronomy three different kinds of time are used, while in ordinary life we use a fourth.

1. Apparent Time—By apparent noon is meant the moment when the sun is on the meridian, and apparent time is measured by the distance in degrees that the sun is east or west of the meridian. Apparent time is given by the sun-dial.

2. Mean Time—The interval between apparent noon on two successive days is not constant, and a clock cannot be constructed to keep apparent time. For this reason mean time is used. The length of a mean day is the average of all the apparent days throughout the year. The real sun moves about the ecliptic in one year; an imaginary mean sun is considered as moving uniformly around the celestial equator in one year. The difference between the times that the real sun and the mean sun cross the meridian is the equation of time. Or, in general, Apparent Time—Mean Time = Equation of Time. This is the same as Correction to Sundial on page 7, with the sign reversed.

3. Sidereal Time—This is time as determined from the stars. It is sidereal noon when the Vernal Equinox or First of Aries is on the meridian. In accurate time-keeping the moment when a star is on the meridian is observed and the corresponding mean time is then computed with the assistance of the Nautical Almanac. When a telescope is mounted equatorially the position of a body in the sky is located by means of the sidereal time.

4. Standard Time—In everyday life we use still another kind of time. A moment's thought will show that in general two places will not have the same mean time; indeed, difference in longitude between two places is determined from their difference in time. But in travelling it is very inconvenient to have the time varying from station to station. For the purpose of facilitating transportation the system of *Standard Time* was introduced in 1883. Within a certain belt approximately 15° wide, all the clocks show the same time, and in passing from one belt to the next the hands of the clock are moved forward or backward one hour.

In Canada we have six standard time belts, as follows;—60th meridian or Atlantic Time, 4h. slower than Greenwich; 75th meridian or Eastern Time, 5h.; 90th meridian or Central Time, 6h.; 105th meridian or Mountain Time, 7h.; 120th meridian or Pacific Time, 8h.; and 135th meridian or Yukon Time, 9h. slower than Greenwich.

The boundaries of the time belts are shown on the map on page 9.

MAP OF STANDARD TIME ZONES



TIMES OF SUNRISE AND SUNSET

In the tables on pages 11 to 16 are given the times of sunrise and sunset for places in latitudes 36° , 40° , 44° , 46° , 48° , 50° and 52° . The times are given in Local Mean Time, and in the table below are given corrections to change from Local Mean to Standard Time for the cities and towns named.

How the Tables are Constructed

The time of sunrise and sunset at a given place, in local mean time, varies from day to day, and depends principally upon the declination of the sun. Variations in the equation of time, the apparent diameter of the sun and atmospheric refraction at the points of sunrise and sunset also affect the final result. These quantities, as well as the solar declination, do not have precisely the same values on corresponding days from year to year, and so the table gives only approximately average values. The times are for the rising and setting of the upper limb of the sun, and are corrected for refraction. It must also be remembered that these times are computed for the sea horizon, which is only approximately realised on land surfaces, and is generally widely departed from in hilly and mountainous localities. The greater or less elevation of the point of view above the ground must also be considered, to get exact results.

The Standard Times for Any Station

In order to find the time of sunrise and sunset for any place on any day, first from the list below find the approximate latitude of the place and the correction, in minutes, which follows the name. Then find in the monthly table the local time of sunrise and sunset for the proper latitude, on the desired day, and apply the correction to get the Standard Time.

		1		1		1	
34°	min.	44°	min.	46°	min.	50°	min
Los Angeles	- 7	Brantford	+21	Glace Bay	0	Brandon	+40
		Guelph	+21	Moncton	+19	Kenora	+18
38°		Halifax	+14	Montreal	~ 6	Medicine Hat	+22
St. Louis	+1	Hamilton	+20	New Glasgow	+11	Moose Jaw	+ 2
San Francisco	+10	Kingston	+ 6	North Bay	+18	Port. la Prairie	+33
Washington	+ 8	Kitchener	+22	Ottawa	+ 3	Regina	- 2
		Milwaukee	- 8	Parry Sound	+20	Trail	- 9
40°		Minneapolis	+13	Quebec	-15	Vancouver	+12
Baltimore	+ 6	Orillia	+18	St. John, N.B.	+24	Winnipeg	+28
New York	- 4	Oshawa	+15	Sault St. Marie			
Philadelphia	+ 1	Owen Sound	+24	Sherbrooke	-12	52°	
Pittsburgh	+20	Peterborough	+13	Sudbury	+24	Calgary	+36
		St. Catharines	+17	Sydney	+ 1	Saskatoon	+ 6
42°		Stratford	+24	Three Rivers	-10		
Boston	-16	Toronto	+18			54°	
Buffalo	+15	Woodstock,Ont		48°		Edmonton	+34
Chicago	-10	Yarmouth	+24	Port Arthur	+57	Prince Albert	+ 1
Cleveland	+26			St. John's, Nfd.		Prince Rupert	+41
Detroit	-28	46°		Seattle	+ 9		
London, Ont.	+25	Charlottetown	+13	Timmins	+26	60°	
Windsor	+32	Fredericton	+26	Victoria	+13	Dawson	+18
							-

Example.—Find the time of sunrise at Owen Sound, also at Regina, on February 11.

In the above list Owen Sound is under " 48° ", and the correction is + 24 min. On page 11 the time of sunrise on February 12 for latitude 44° is 7.09; add 24 min. and we get 7.33 (Eastern Standard Time). Regina is under "50°", and the correction is -2 min. From the table the time is 7.17 and subtracting 2 min. we get the time of sunrise 7.15 (Mountain Standard Time).

DATE Latitude 40° Latitude 50° Latitu																	
DATE Suntise Sunset Suntise Sunset <th></th> <th></th> <th>La</th> <th>titud</th> <th>e 36°</th> <th>Latitu</th> <th>de 40°</th> <th>Latitu</th> <th>de 44°</th> <th>Latitı</th> <th>ide 46°</th> <th>Latitu</th> <th>ide 48°</th> <th>Latitu</th> <th>de 50°</th> <th>Latitu</th> <th>de 52</th>			La	titud	e 36°	Latitu	de 40 °	Latitu	de 44 °	Latitı	ide 46 °	Latitu	ide 48 °	Latitu	de 50 °	Latitu	de 5 2
January 7 7 111 457 723 447 735 432 756 413 756 413 863 7 7111 500 722 447 735 433 742 426 757 418 806 441 7 7111 500 722 456 735 438 742 426 757 418 806 441 7 7111 500 722 456 733 446 741 433 756 420 757 418 806 117 710 510 722 455 733 444 744 433 756 423 806 443 117 710 512 720 456 733 444 744 433 756 423 806 443 117 710 512 720 456 733 444 744 433 756 423 117 710 512 710 512 720 453 733 444 744 433 756 444 117 706 512 716 512 716 512 716 423 756 423 756 444 117 710 512 716 513 722 500 733 444 744 746 444 746 444 746 206 523 710 523 710 523 710 <t< th=""><th>DATE</th><th>×</th><th>Sun</th><th>rise \$</th><th>Sunset</th><th>Sunrise</th><th>Sunset</th><th>Sunrise</th><th></th><th>Sunrise</th><th></th><th>Sunrise</th><th></th><th>Sunrise</th><th></th><th>Sunrise</th><th></th></t<>	DATE	×	Sun	rise \$	Sunset	Sunrise	Sunset	Sunrise		Sunrise		Sunrise		Sunrise		Sunrise	
8 7 11 56 7 55 447 755 447 755 447 755 447 755 447 755 447 755 447 755 447 755 447 755 447 757 418 805 447 757 418 755 448 756 448 756 448 756 448 756 448 756 448 756 448 756 448 756 448 756 448 756 448 756 448	January	-	чЪ														1
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		ю.	~														
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DATE		Latitu Sunrise	Latitude 36° Sunrise Sunset	Latitude Sunrise Sur	ide 40 ° sunset	Latitude 44 ° Sunrise Sunset	Latitude 46 ° Sunrise Sunset	Latitude 48 ° Sunrise Sunset	Latitude 50 ° Sunrise Sunset	Latitude 52° Sunrise Sunset
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DATE		Latitu Sunrise	Latitude 36° Sunrise Sunset	Latitude 40 ° Sunrise Sunset	Latitud e 44 ° Sunrise Sunset	Latitude 46 ° Sunrise Sunset	Latitude 48 ° Sunrise Sunset	Lat	Latitude 50 ° Sunrise Sunset
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November 1 $\begin{array}{cccccccccccccccccccccccccccccccccccc$	DATE	Lati Sunri	Latitude 36° Sunrise Sunset	Latitud e 40 ° Sunrise Sunset	d e 40 ° Sunset	Latitude 44 ° Sunrise Sunset	d e 44 ° Sunset	Latitu Sunrise	Latitude 46 ° Sunrise Sunset	Latitude Sunrise Sun	ide 48 ° Sunset	Latitude Sunrise Sun	d e 50 ° Sunset	Latitude Sunrise Su	ide 52 ° Sunset
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THE SUN AND PLANETS FOR 1939

By DONALD A. MACRAE

THE SUN

It is a well-known fact that the variations in the number and positions of sun-spots observed on the sun are roughly periodic. The average interval between times of maximum solar activity is 11.13 years, but the observed intervals range from eight to fourteen years. The rise to maximum is usually more rapid than the fall which follows. Since 1934 solar activity has been increasing and a maximum was expected early in 1938. Minor fluctuations since August 1937 have made it difficult to decide whether the sun has already passed the maximum or is still in a state of increasing activity. In either case sun-spots will be numerous and the year 1939 will offer an abundance of magnetic storms, auroral displays and other associated phenomena.

MERCURY

Mercury, the planet closest to the sun, is also the smallest and least massive. With the exception of Pluto its orbit is the most eccentric and has the greatest inclination to the ecliptic. Mercury appears to move swiftly from one side of the sun to the other several times each year and at times of greatest elongation its angular distance from the sun is always small, varying from 18° to 28°. It is visible to the naked eye for about two weeks at these times.

When Mercury is near greatest elongation east of the sun it appears in the evening, setting very soon after the sun. When near greatest western elongation it can be seen in the morning just before sunrise. In northern and southern latitudes at sunset the ecliptic is most nearly vertical in the spring; at sunrise it is most nearly vertical in the autumn. Therefore eastern elongations in the spring and western elongations in the autumn are most favourable for observing Mercury.

Mercury reaches eastern elongation three times during 1939. The dates, angular distances from the sun, and magnitudes are: March 16 (most favour-able), 18°, +0.0; July 13, 27°, +0.7; November 7, 23°, +0.0.

Mercury reaches western elongation four times as follows, January 3, 23°, +0.0; May 1, 27°, +0.7; August 28 (most favourable), 18°, +0.1; December 16, 21°, -0.2.

At its closest approach to the earth this year (inferior conjunction, April 3) its distance will be 54,000,000 miles. At greatest elongations its semidiameter is between three and four seconds of arc.

VENUS

Venus is the next planet in order from the sun. In size and mass and perhaps in other respects it resembles the earth. To us it appears as the most brilliant "star" in the sky. Venus performs in the same way as Mercury but moves much more slowly and is farther removed from the sun. The time for one complete oscillation is 1.6 years and greatest elongation is about 45° . When west of the sun Venus appears as the morning star and in this position it was known to the ancients as Phosphorus, the light-bringer. When east of the sun it is the evening star, Hesperus.

At the beginning of the year Venus is west of the sun and at its greatest brilliancy, magnitude -4.4, thirteen times as bright as Sirius. Greatest western elongation (47° and magnitude -4.1) occurs on January 30. January, in fact, is the most favourable time for observation; after this month Venus moves in toward the sun again very slowly but it rapidly becomes more poorly situated for observation because of the position of the ecliptic at sunrise. Superior conjunction occurs on September 5, and from then until the end of the year it will be moving eastward away from the sun, and will set soon after sunset.

After April its magnitude will remain close to -3.4. Since Venus is on the far side of the sun most of the year it is never very close to the earth and its semidiameter decreases quickly from 18 to about 5 seconds of arc.

MARS

Mars is the fourth planet from the sun and the first superior planet. Its path in the sky is similar to all planets beyond the earth, a slow motion in the region of the zodiac from west to east with occasional periods when its motion is retrograde. Mars is best observed at favourable oppositions, when the planet is near its perihelion at the time of opposition. These occur every 15 or 17 years, always in the latter part of August, and the next one will be in 1941. Nevertheless Mars is very well placed for observation during 1939 and at the time of opposition will be within 36,000,000 miles of the earth, as compared with 34,600,000 at the most favourable oppositions.

On January 1 it is of magnitude +1.7, four hours west of the sun in the constellation Libra, and 180 million miles from the earth. It increases in brightness and remains in the morning sky for the first half of the year. From June 24 to August 24 its motion is retrograde in the constellation Capricornus, and opposition comes on July 23, when it is of magnitude -2.6, 2.5 times as bright as Sirius. Its diameter is then about 24 seconds of arc. During the latter part of the year it is an evening star of magnitude about +0.0 in the constellations Capricornus and Aquarius. Eastern quadrature occurs on November 29. The accompanying chart gives its path among the stars during the year.





THE ASTEROIDS

Between the orbits of Mars and Jupiter there are a large number of small bodies revolving about the sun. The first of these minor planets was discovered in 1801, and the number with orbits determined is now about 1400. The majority are less than 50 miles in diameter. They all revolve from west to east, and some approach very close to the earth.

In most telescopes these asteroids show no discs but because of their swift motions among the stars they make very interesting subjects for observation. Ephemerides of the brighter asteroids are found on page 24.

JUPITER

Jupiter is the largest and the most massive planet of the solar system. Because of its distance from the sun and the earth, its motion among the stars is quite slow. An evening star of magnitude -1.6 at the beginning of the year,



conjunction with the sun occurs on March 6. After this it will be visible in the early morning sky, western quadrature coming on June 29. From July 30 to November 25 its motion is retrograde. Opposition occurs on September 27, when it is of maximum brightness for the year, magnitude -2.5. It is in eastern quadrature on December 22 still a brilliant object at magnitude -2.0. For its path among the stars, see the accompanying chart.

Jupiter's period of rotation is the shortest of all the planets, about ten hours; as a result there is a marked flattening at the poles. Four of its moons can be seen with a good pair of binoculars. Their configurations are given elsewhere. Seven other moons (two discovered in 1938) are so faint they are visible in only the largest telescopes. The moons, the surface detail, its large disc and its position in the sky make Jupiter a very interesting object for observation. In September its apparent semidiameter is about 23" and its distance from the earth is 367 million miles.

SATURN

Saturn is the next planet in order from the sun. It is also next to Jupiter in size and mass. Its motion in the heavens is very slow. It remains about two hours east of Jupiter all year and as a result its configurations except that they occur about a month later are similar to Jupiter's. Saturn begins the year in eastern quadrature (January 2) with magnitude +0.9. Conjunction with the sun occurs on April 11. In the summer it will be visible in the morning sky, western quadrature being on July 24, when it is of magnitude +0.6. After



August 14 its motion is retrograde. At opposition on October 21 its magnitude is +0.2 and it will be visible all night as a yellowish star. Toward the end of the year it will be an object in the evening sky, at magnitude about +0.5. In October 1939 its distance from the earth is 770 million miles and its semidiameter is almost 9".

Saturn's unique ring system makes it one of the most interesting objects in our skies. These rings, the outer ring, the bright ring, and the crape ring, are composed of a large number of very small satellites which revolve about Saturn in a plane inclined to the planet's orbit at an angle of 27°. In 1936 the rings were presented edge on and so were invisible. They will appear opened out to their maximum extent in 1943, when the planet will be in an excellent position for observation in the northern hemisphere.

URANUS

The ancient astronomers were well familiar with the first six planets. The seventh, Uranus, was not discovered until telescopic observation was firmly established. To Sir William Herschel goes the credit for finding this body, which he at first thought was a comet. Later observations proved it to be the next planet beyond Saturn. Herschel suggested calling it *Georgium Sidus* after George III. During 1939 it will appear as a blue-green sixth magnitude star about one and a half hours east of Saturn in the constellation Aries. Its semi-diameter is 1.8 seconds of arc.



Eastern quadrature is on February 3 so that for the first three months it can be observed in the evening. From March to June it is too near the sun for good observation. Conjunction occurs on May 9. Western quadrature is on August 15 and opposition on November 13, when it is above the horizon all night. The path of Uranus among the stars is given in the chart. Its motion is direct from January 22 to August 28.

NEPTUNE

Although Uranus was discovered by accident, the next planet was found by means of the so-called "astronomy of the invisible". The story of its almost simultaneous discovery in 1846 by Leverrier and Adams is one of the most interesting in Astronomy. The observed deviations of Uranus from its calculated orbit led them both to predict correctly the position in the sky of the perturbing planet Neptune. In 1939 Neptune appears as a blue-green eighth magnitude star in the constellations Leo and Virgo. It is conveniently situated for observation in the first half of the year, reaching opposition on March 13 at magnitude 7.7, and eastern quadrature on June 12. From August to October it is rather close to the sun, passing conjunction on September 16. It is in western quadrature on December 18, and can be seen in the morning. The accompanying chart will identify Neptune among the stars; until June 2 the planet is retrograding.

Neptune's rotation period is quite short. It has been determined spectrographically as 1534 hours. At opposition it is about 2,700 million miles from the earth and has an apparent semidiameter of 1.25 seconds of arc.



PLUTO

The success of the theory of perturbations in this field led Lowell to investigate the existence of a trans-Neptunian planet. The observatory which he founded announced the discovery of Pluto reasonably near its predicted position on March 13, 1930, the anniversary of Lowell's birth and of Herschel's discovery of Uranus.

During 1939 Pluto is a yellowish star in the constellation Cancer, just south of λ Cancri. It is about magnitude 15 and so is invisible in all but the largest telescopes. The position, which changes only slightly during the year, on August 1 is

 $a : 8^{h} 17^{m}.9$ $\delta : +23^{\circ} 10'$

OPPOSITION EPHEMERIDES OF THE BRIGHTEST ASTEROIDS, 1939.

	1	29 Amphitrite	
$\begin{array}{cccccccccccccccccccccccccccccccccccc$		$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{r} + 30 & 43 \\ + 30 & 59 \\ + 31 & 07 \\ + 31 & 05 \end{array}$
Opp. Aug. 19 Mag. 8	8.3	Opp. Jan. 13	Mag. 8.8
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$2 42' \\ 00 \\ 22 \\ 52 \\ 31$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$
	22 6.9	1623 45.0	- 9 44 Mag. 8.9 ³
$\begin{array}{c} 7 \ \text{Iris} \\ \text{June 1819 } ^{h}24.8^{m}-18^{\circ} \\ 2619 \ 17.7 \ -18 \\ \text{July } 419 \ 09.6 \ -18 \\ 1219 \ 00.9 \ -18 \\ 2018 \ 52.4 \ -18 \end{array}$	45 41 37 35	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$
Opp. July 7 Mag.	8.7	Opp. Apr. 13	Mag. 8.6
$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	$\begin{array}{c c} 31 \\ 32 \\ 28 \\ 16 \end{array}$	1817 27.0	$\begin{array}{rrrrr} - & 24 & 20 \\ - & 23 & 51 \\ - & 23 & 18 \\ - & 22 & 45 \end{array}$
Opp. Aug. 30 Mag.	8.6	Opp. June 14	Mag. 8.8
$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	47	511 DAVIDA Oct. 24	$n - 4^{\circ} 23' - 4 47 - 5 00 - 4 59 - 4 44 - 4 14$
Opp. Aug. 5 Mag.	8.1	Opp. Nov. 13	Mag. 8.7

PREPARED BY PROFESSOR G. STRACKE

These ephemerides have been made available through the kindness of Professor Stracke and of Professor A. Kopff, the Director of the Astronomisches Rechen-Institut, of Berlin. Positions are for equinox of 1950.

ECLIPSES, 1939

In 1939 there will be four eclipses, two of the sun and two of the moon.

I. An Annular Eclipse of the Sun, 1939 April 19, visible in Canada. The path of annular eclipse begins in the Aleutian Islands, passes through Alaska, over the mouth of the Mackenzie River, and ends after passing over the North Pole. The eclipse will be visible in its partial phases throughout the whole of North America and the North Atlantic. In western Canada, the eclipse will begin about 10^{h} and end about 12^{h} 30^{m} . In eastern Canada it will begin about 11^{h} and end about 13^{h} (75th Meridian Civil Time).

II. A Total Eclipse of the Moon, 1939 May 3, invisible in Canada. Visible only in Asia, Africa, and part of Europe.

III. A Total Eclipse of the Sun, 1939 October 12, invisible in Canada. The path of totality is short and passes close to the South Pole. The duration of the total phase is about one and one-half minutes. The eclipse is visible in its partial phase in the southern Pacific Ocean.

IV. A Partial Eclipse of the Moon, 1939 October 27, visible in Canada. The beginning is visible generally in Europe except the extreme eastern part, the western part of Africa, the Atlantic Ocean, North and South America, the eastern part of the Pacific Ocean, and the north-eastern extremity of Asia. The ending is visible generally in the North Atlantic Ocean, the Arctic Ocean, North and South America, the Pacific Ocean, Polynesia, the eastern part of Australia, and north-eastern Asia.

Circumstances of the Eclipse (75th Meridian Civil Time)

	d	h	m
Moon enters penumbraOctober	27	22	42
Moon enters umbra "	27	23	54
Middle of eclipse "	28	01	3 6
Moon leaves umbra "	28	03	18
Moon leaves penumbra "	28	04	31

The Magnitude of the eclipse is 0.992 (moon's diameter=1.0). First contact of the umbra with the moon's limb occurs 48° east of the north point and last contact 77° west of the north point.

THE SKY MONTH BY MONTH

By P. M. Millman

THE SKY FOR JANUARY, 1939

The times of transit are given in local mean time, 0h at midnight; to change to Standard Time, see p. 10. Estimates of altitude are for an observer in latitude 45° N.

The Sun—During January the sun's R.A. increases from 18h 42m to 20h 55m, and its Decl. changes from 23° 6'S. to $17^{\circ} 25'$ S. The equation of time (see p. 7) decreases from -3m 6s to -13m 33s. Owing to this rapid drop in value the time of mean noon appears, for the first ten days of the month, to remain at the same distance from sunrise, that is, the forenoons as indicated by our clocks are of the same length. On the 20th of the month the sun moves into the sign Aquarius, the second winter zodiacal sign. The signs of the zodiac are all exactly 30° in length, commencing at the first point of Aries. Though the sign of Aquarius once corresponded to the constellation of the same name, it now is situated among the stars of Capricornus, owing to the westward motion of the first point of Aries. The earth is nearest the sun, that is in perihelion, on January 3.

The Moon—For its phases, perigee and apogee times and distances, and its conjunctions with the planets, see opposite page. The moon occults the planet Uranus on the 29th of the month.

Mercury on the 15th is in R.A. 18h 16m, Decl. 23° 14' S. and transits at 10.43. Mercury reaches its greatest elongation west of the sun in the morning sky on the 3rd and for a few days before and after this date will be well placed for observation rising nearly 2 hours before the sun in the S.E. and being about 19° above the horizon at sunrise. It is a reddish star of magnitude 0.

Venus on the 15th is in R.A. 16h 28m, Decl. 17° 33' S. and transits at 8.54. It is a brilliant object in the morning sky rising over 3 hours before the sun, Greatest elongation west is on the 30th.

Mars on the 15th is in R.A. 15h 16m, Decl. 17° 14' S. and transits at 7.41. It is slowly moving into the morning sky, being in view most of the last half of the night. It is in close conjunction with the moon on the 14th.

Jupiter on the 15th is in R.A. 22h 22m, Decl. $11^{\circ} 14'$ S. and transits at 14.46. It is a bright star in the evening sky setting over 3 hours after the sun. For the configurations of its satellites see next page and for their eclipses, etc. see p. 50.

Saturn on the 15th is in R.A. 0h 48m, Decl. $2^{\circ} 32'$ N. and transits at 17.11. It is in quadrature with the sun on the 2nd and appears just east of the meridian at sunset, remaining in view for the first half of the night. For the elongations, etc., of its satellites see p. 52.

Uranus on the 15th is in R.A. 2h 46m, Decl. 15° 40' N. and transits at 19.08. Neptune on the 15th is in R.A. 11h 37m, Decl. 3° 46' N. and transits at 4.01. Pluto—For information regarding this planet, see p. 23.

ASTRONOMICAL PHENOMENA MONTH BY MONTH

				JANUARY	Config.
				Min.	of Jupiter's
				75th Meridian Civil Time of Algol	Sat. 18h 45m
	d	h	m	h m	
Sun.	1	13	09	ර් හී ଐ රී 0° 35′ S	32104
Mon.		15		$\Box \flat \odot$ Quadrantid Meteors. p. 54	30124
Tue.	3	5		§ Greatest elongation W., 22° 49'	31024
		17		\oplus in Perihelion. Dist. from \odot , 91,340,000 mi.	
Wed.	4	10		\bigcirc in Perihelion	d 2 O13
Thu.	5	16	30	⁽²⁾ Full Moon	42103
Fri.	6	6		Moon in Perigee. Dist. from \oplus , 221,900 mi	d4O23
Sat.	7				42031
Sun.	8				43210
Mon.	9			••••••••••••••••	43021
Tue.	10	10	22	$\sigma' \Psi \mathbb{G} \qquad \Psi \qquad 5^{\circ} 31' \mathrm{N}$	43102
Wed.	11				42031
Thu.	12	8	10	C Last Quarter	24103
Fri.	13				01423
Sat.	14	16	00	ර් ර් ℃	dO34*
Sun.	15				23104
Mon.	16	1		\$ in \$\constraint\$	30214
		4	12	o´♀	
Tue.	17			••••••••••••••••••	31024
Wed.	18	15	25	σ′₿ (ξ 4° 25′ S	20314
Thu.	19				21034
Fri.	20	8	27	New Moon	01243
		18		Moon in Apogee. Dist. from \oplus , 252,700 mi	
Sat.	21				dO23*
Sun.	22	10		Stationary in R.A14	42310
Mon.	23	12	4 4	σ′21 C 24 6° 00′ S	4301*
Tue.	24				43102
Wed.	25				4201*
Thu.	26	7		Q Greatest Hel. Lat. N	42103
		8		§ in Aphelion.	
		14	44	♂ þ € þ 5° 13′ S.	
Fri.	27			· · · · · · · · · · · · · · · · · · ·	40123
Sat.	28	10	00	b First Quarter	41023
		21	15	ີດ້ ອີ ⊈ີ້ ອີ 0° 20′ S.	
Sun.	29				d234O
Mon.	30	8		\bigcirc Greatest elongation W., 46° 56′	32014
Tue.	31				31024

By Ruth J. Northcott

Explanation of symbols and abbreviations on p. 4, of time on p. 8.

THE SKY FOR FEBRUARY, 1939

The times of transit are given in local mean time, 0h at midnight; to change to Standard Time, see p. 10. Estimates of altitude are for an observer in latitude 45° N.

The Sun—During February the sun's R.A. increases from 20h 55m to 22h 44m, and its Decl. changes from $17^{\circ} 25'$ S. to $8^{\circ} 0'$ S. The equation of time decreases from -13m 33s to a minimum of -14m 21s. on Feb. 12, and then rises to -12m 42s at the end of the month (see p. 7). For changes in the length of the day see p. 11. The sun enters the sign Pisces, the third winter sign, on the 19th.

The Moon—For its phases, perigee and apogee times and distances, and its conjunctions with the planets, see opposite page.

Mercury on the 15th is in RA. 21h 41m, Decl. $16^{\circ} 5'$ S. and transits at 12.07. It is in superior conjunction with the sun on the 18th and too near the sun for observation during February.

Venus on the 15th is in R.A. 18h 39m, Decl. $20^{\circ} 30'$ S. and transits at 9.03. It is a brilliant yellow star of magnitude -4, rising in the south-east almost three hours before the sun.

Mars on the 15th is in R.A. 16h 34m, Decl. $21^{\circ} 26'$ S. and transits at 6.57. It is a red star of the first magnitude in the morning sky, rising about an hour after midnight.

Jupiter on the 15th is in R.A. 22h 49m, Decl. 8° 36' S. and transits at 13.10. It is fast approaching the sun in the evening sky, setting a little over an hour after sunset. For the configurations of its satellites see next page and for their eclipses, etc., see p. 50.

Saturn on the 15th is in R.A. 0h 57m, Decl. $3^{\circ} 34'$ N. and transits at 15.18. It is a pale yellow star slightly brighter than first magnitude and sets about four hours after the sun in the evening sky. For the elongations, etc., of its satellites see p. 52.

Uranus on the 15th is in R.A. 2h 47m, Decl. 15° 44' N. and transits at 17.07. *Neptune* on the 15th is in R.A. 11h 35m, Decl. 4° 0' N. and transits at 1.57. *Pluto*—For information regarding this planet, see p. 23.

				FEBRUARY	Min.	Config.
				75th Meridian Civil Time	of Algol	Jupiter's Sat. 18h 30m
	d	h	m		h m	
Wed.	1			· · · · · · · · · · · · · · · · · · ·	• •	32014
Thu.	2					21034
Fri.	3	10		□ ô ⊙	1 33	01234
		19		Moon in Perigee. Dist. from \oplus , 221,600 mi.		
Sat.	4	2	55	Full Moon		10234
Sun.	5				. 22 23	23014
Mon.	6	19	09			3204*
Tue.	7					314O2
Wed.	8				19 12	d43O1
Thu.	9					42103
Fri.	10	23	12	C Last Quarter		40213
Sat.	11				. 16 01	41023
Sun.	12	7	31	ଏ ଟି ଐ ସି 1° 40′ S		d42O1
Mon.	13					43210
Tue.	14	21	47	ଏହୁ ⊈ ଦୁ 1° 36′ S	.12 51	d34O2
Wed.	15	16		g Greatest Hel. Lat. S		34021
Thu.	16	21		Moon in Apogee. Dist. from \oplus , 252,600 mi	••	21034
Fri.	17				. 9 40	O2134
Sat.	18	21				10234
Sun.	19	3	28	New Moon.		
		9	54	σ'₿ Œ ₿ 7° 02' S.		
Mon.	20	7	25	σ 24 € 24 5° 38′ S	. 6 29	
Tue.	21					
Wed.	22					
Thu.	23	0	43	♂ þ € þ 4° 41′ S	. 3 19	
Fri.	24					
Sat.	25	4	11	රô⊈ ô 0°00′.		
Sun.	26	0		σ ['] [†] ² ⁴ [†] ² ¹ [†] ² ¹ [†] ¹ ^{†[†]¹[†]¹^{†[†]¹^{†[†][†]^{†[†]^{†[†][†]^{†^{†[†]^{†[†]^{†^{†[†]^{†^{†[†]^{†^{†[†]^{†^{†[†]^{†^{†[†]^{†^{†[†]^{†^{†[†]^{†^{†[†]^{†^{†^{†[†]^{†^{†[†]^{†^{†[†]^{†^{†[†]^{†^{†^{†[†]^{†^{†^{†[†]^{†^{†^{†^{†^{†[†]^{†^{†^{†^{†^{†^{†^{†^{†^{†^{†^{†^{†^{†^{†^{†^{†^{†^{†^{†[†]}}}	. 0 08	
		22	26			
Mon.	27			~ ~		
Tue.	28			•••••••••••••••••••••••••••••••••••••••	.20 57	

Explanation of symbols and abbreviations on p. 4, of time on p. 8. Jupiter being near the Sun, phenomena of the satellites are not given from February 19 to May 1.

THE SKY FOR MARCH, 1939

The times of transit are given in local mean time, 0h at midnight; to change to Standard Time, see p. 10. Estimates of altitude are for an observer in latitude 45° N.

The Sun—During March the sun's R.A. increases from 22h 44m to 0h 38m and its Decl. changes from $8^{\circ} 0'$ S. to $4^{\circ} 7'$ N. The equation of time increases from -12m 42s to -4m 19s (see p. 7). For changes in the length of the day, see p. 12. The sun is at the vernal equinox and crosses the equator on its trip north at 12h 29m G.C.T. on the 21st. This date marks the beginning of spring and day and night are approximately equal all over the world.

The Moon—For its phases, perigee and apogee times and distances, and for its conjunctions with the planets, see opposite page.

Mercury on the 15th is in R.A. 0h 40m, Decl. 6° 13' N. and transits at 13.13. It is at greatest elongation east of the sun on the 16th. At this time it appears in the evening sky being about 18° above the horizon at sunset and setting approximately 1h 40m after the sun. This is the most favourable elongation of the year for observing Mercury in the evening sky.

Venus on the 15th is in R.A. 20h 53m, Decl. $16^{\circ} 57'$ S. and transits at 9.26. It is slowly growing fainter as it recedes from the earth but is still very prominent in the morning sky, rising just under 2 hours before the sun.

Mars on the 15th is in R.A. 17h 44m, Decl. 23° 18' S. and transits at 6.17. It is in quadrature on the 21st but owing to its southern declination it does not rise till some time after midnight.

Jupiter on the 15th is in R.A. 23h 14m, Decl. 6° 1' S. and transits at 11.45. It is in conjunction with the sun on March. 6 and too near that body to be observed this month.

Saturn on the 15th is in R.A. 1h 8m, Decl. 4° 48' N. and transits at 13.39. It is in the S.W. at sunset and sets about 2 hours after the sun. It is rapidly approaching the sun in the evening sky and by the end of the month will be unfavourably placed for observation. For the elongations, etc., of its satellites, see p. 52.

Uranus on the 15th is in R.A. 2h 50m, Decl. 16° 0' N. and transits at 15.20.

Neptune on the 15th is in R.A. 11h 32m, Decl. 4° 18' N. and transits at 0.05. Opposition to the sun is on the 13th.

Pluto-For information regarding this planet, see p. 23.

				MARCH	Min. of
				75th Meridian Civil Time	Algol
	d	h	m		h m
Wed.	1				•••••
Thu.	2			•••••••••••••••••••••••••••••••••••••••	• • • • • •
Fri.	3				
Sat.	4	6		Moon in Perigee. Dist. from \oplus , 223,300 i	ni.
Sun.		13			
Mon.	6	5 7 16		 <i>σ</i>Ψ € Ψ 5° 16′ N <i>σ</i> 24 ⊙ <i>ψ</i> in Ω 	14 36
Tue.	7				
Wed.	8				
Thu.	9				
Fri.	10			•••••••••••••••••••••••••••••••••••••••	
Sat.	11	7		§ in Perihelion.	
Sun.	12	16	37	Last Quarter	8 14
Mon.			16	ഗ്ഗ്് ഗ് 3° 33′ S.	
		6		$\mathcal{O}\Psi \odot$ Dist. from \oplus , 2,715,000,000) mi.
Tue.	14				
Wed.	15				
Thu.	16	10		Moon in Apogee. Dist. from \oplus , 252,100 r	ni.
		20		Greatest elongation E., 18° 27'.	
Fri.		4	29	ơ♀₵ ♀ 4°54′S.	
Sat.	18				$\dots 153$
Sun.		10		o ⁷ in ¹⁰	
Mon.	20			o 24 € 24 5° 17′ S	$\dots 22 42$
~	~ 1		49		0
Tue.	21		29	\odot enters Υ , Spring commences. Long. of	⊙, 0°.
		14		§ Greatest Hel. Lat. N.	
17.1	00	14			
wea.	22			 	
T L	<u></u>		17		10.91
Thu. Fri.	23 24	8		9 in °C	19 31
r f1 .	4 4	2	94	 	
Sat.	25	11	4 4	$\mathbf{O} \circ \mathbf{U} \qquad \mathbf{O} \circ 10 \mathbf{N}.$	
Sun.	25 26				
Jun.					
Mon		7	16	· · · · · · · · · · · · · · · · · ·	• • • • • •
Mon. Fue	28		T O	Ju Inst Quarter.	
Гue.	28 29	•			13 10
	-	'			

Explanation of symbols and abbreviations on p. 4, of time on p. 8. Jupiter being near the Sun, phenomena of the satellites are not given from February 19 to May 1.

THE SKY FOR APRIL, 1939

The times of transit are given in local mean time, 0h at midnight; to convert to Standard Time, see p. 10. Estimates of altitude are for an observer in latitude 45° N.

The Sun—During April the sun's R.A. increases from 0h 38m to 2h 29m and its Decl. changes from 4° 7' N. to 14° 44' N. The equation of time rises from -4m 19s to +2m 46s (see p. 7). For changes in the length of the day, see p. 12. On the 20th the sun enters the sign Taurus, the second spring sign of the zodiac. There is an annular eclipse of the sun on the 19th, visible in Canada and the United States as a partial eclipse. For details see p. 29.

The Moon—For its phases, perigee and apogee times and distances, and its conjunctions with the planets, see opposite page.

Mercury on the 15th is in R.A. 0h 22m, Decl. 2° 7' N. and transits at 10.50. It is in inferior conjunction with the sun on the 3rd and too near that body for observation during most of the month. During the last few days of the month it is approaching western elongation and will be visible in the morning sky.

Venus on the 15th is in R.A. 23h 15m, Decl. 6° 5' S. and transits at 9.47. It is a bright star in the morning sky, rising about an hour and a half before the sun. On April 21st it is in close conjunction with Jupiter, the two planets being less than half a degree apart. For a few days Venus and Jupiter will form a fine double star in the morning sky.

Mars on the 15th is in R.A. 18h 58m, Decl. $23^{\circ} 28'$ S. and transits at 5.28. It is steadily growing brighter as opposition is approached, being now a red star of 0 magnitude. It is in view during the last half of the night.

Jupiter on the 15th is in R.A. 23h 41m, Decl. 3° 12' S. and transits at 10.10. It is near the sun in the morning sky and not well placed for observation. On the 21st it is in close conjunction with Venus.

Saturn on the 15th is in R.A. 1h 23m, Decl. 6° 16' N. and transits at 11.51. It is in conjunction with the sun on the 11th and too near that body to be observed during this month.

Uranus on the 15th is in R.A. 2h 56m, Decl. 16° 27' N. and transits at 13.24. Neptune on the 15th is in R.A. 11h 29m, Decl. 4° 37' N. and transits at 21.56. Pluto—For information regarding this planet, see p. 23.

Normalized and the image of the image o	APRIL Min.	
Sat. 1 8 Moon in Perigee. Dist. from \bigoplus , 226,500 mi 59 Sun. 2 14 08 $\mathcal{O} \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \ \$		
Sun. 2 14 08 $\sigma \ \Psi \ \Theta \ \Theta$		
Mon. 3 3 $\sigma' \buildreft \ \odot$ Inferior. 23 18 Full Moon. 6 48 Wed. 5 6 48 Wed. 5 6 48 Fri. 7 6 3.37 Sat. 8 6 3.37 Sun. 9 7 5° 16' S. 0 Mon. 10 16 33 $\sigma' \sigma' \buildreft \b$		
23 18 Full Moon. Tue. 4 6 Wed. 5 6 Thu. 6 6 Fri. 7 3 Sat. 8 3 Sun. 9 3 Mon. 10 16 33 $\sigma' \sigma' \oplus \sigma'' \oplus \sigma'' \circ 16' S$		
Wed. 5 3 37 Thu. 6 3 37 Fri. 7 3 37 Sat. 8 3 37 Mon. 10 16 33 $\sigma' \sigma^3 \oplus \oplus \oplus \sigma^3 \oplus \oplus \oplus \oplus \sigma^3 \oplus \oplus$		
Thu. 6		
Fri. 7	• • • • • • • • • • • • • • • • • • • •	
Sat. 8 Sun. 9 Mon. 10 16 33 $\sigma' \ \ensuremath{\mathbb{C}}$ 5° 16' S	•••••••••••••••••	
Sun. 9	3 37	
Mon. 10 16 33 $\sigma ? \blacksquare = \sigma^3 = 5^\circ 16' \text{ S}$		
Tue. 11 11 11 I Last Quarter. 15 $\sigma' \flat \odot$		
15 $\sigma \not b \odot$ Wed. 12 Thu. 13 4 Moon in Apogee. Dist from \bigoplus , 251,400 mi. Fri. 14 1 § in °° Sat. 15 13 § Stationary in R.A		
Wed. 12	2	
Thu. 13 4 Moon in Apogee. Dist from \bigoplus , 251,400 mi. Fri. 14 1 \emptyset in \heartsuit Sat. 15 13 \emptyset Stationary in R.A		
Fri.141 \S in \Im Sat.1513 \S Stationary in R.A		
Sat. 15 13 $\ensuremath{\emptyset}$ Stationary in R.A		
Sun. 16 12 $54 \circ \circ \varphi \oplus \varphi$ $5^{\circ} 32' S.$ 23 04 $\circ \circ 24 \oplus \varphi$ $24 + 4^{\circ} 57' S.$ Mon. 17 19 09 $\circ \circ \varphi \oplus \varphi$ ξ $3^{\circ} 59' S.$		
$\begin{array}{cccccccccccccccccccccccccccccccccccc$		
Mon. 17 19 09 ♂ ⊈ (월 3° 59′ S.	ΥΨ ¥ 0°32'S. /0.	
Wed. 19 Annular Eclipse of \bigcirc , see p. 25.		
1 41 σ b \square b 3° 53′ S.		
11 35 (b) New Moon.		
Thu. 20 20 22 0 8 0 8 0° 27' N.		
Fri. 21 21 $\sigma' \neq 24$ φ $0^{\circ} 24' S11 43$		
Lyrid Meteors, p. 54		
Sat. 22		
Sun. 23		
Mon. 24 7 Ø in Aphelion	in Aphelion	
Tue. 25	•••••••••••••••••••	
Wed. 26 13 25 🕽 First Quarter.		
17 \bigcirc in Aphelion.	in Aphelion.	
Thu. 27		
Fri. 28 5 Moon in Perigee. Dist. from \oplus , 229,400 mi.		
Sat. 29 21 06 $\sigma' \Psi \mathbb{G}$ Ψ 5° 24' N.		
Sun. 30		

Explanation of symbols and abbreviations on p. 4, of time on p. 8. Jupiter being near the Sun, phenomena of the satellites are not given from February 19 to May 1.

THE SKY FOR MAY, 1939

The times of transit are given in local mean time, 0h at midnight. To convert to Standard Time, see p. 10. Estimates of altitude are for an observer in latitude 45° N.

The Sun—During May the sun's R.A. increases from $2h \ 29m$ to $4h \ 32m$ and its Decl. changes from $14^{\circ} \ 44'$ N. to $21^{\circ} \ 54'$ N. The equation of time increases from $+2m \ 46s$ to a maximum of $+3m \ 45s$ on the 15th and then drops to $+2m \ 30s$. at the end of the month (see p. 7). For changes in the length of the day see p. 13. On May 21 the sun enters Gemini, the third spring sign of the zodiac.

The Moon—For its phases, perigee and apogee times and distances, and its conjunctions with the planets, see opposite page. There is a total eclipse of the moon on May 3, invisible in Canada and the United States. For details see p. 29.

Mercury on the 15th is in R.A. 1h 58m, Decl. 9° 9' N. and transits at 10.32. On May 1 it reaches greatest elongation west of the sun but is not particularly well placed for observation, rising barely an hour before the sun and being about 9° above the horizon at sunrise. It is of magnitude 0.7. On May 10 it is in close conjunction with Saturn.

Venus on the 15th is in R.A. 1h 29m, Decl. 7° 25' N. and transits at 10.02. Its magnitude has now dropped to -3.4 and it rises just over an hour before the sun.

Mars on the 15th is in R.A. 19h 56m, Decl. $22^{\circ} 42'$ S. and transits at 4.28. It is gradually coming into the evening sky, rising about midnight in the S.E. It has brightened to magnitude -0.6.

Jupiter on the 15th is in R.A. 0h 4m, Decl. 0° 46' S. and transits at 8.35. It is a bright star, rising almost due east about 2 hours before the sun. For the configurations of its satellites see next page, and for their eclipses, etc., see p. 50.

Saturn on the 15th is in R.A. 1h 36m, Decl. 7° 36' N. and transits at 10.07. It is in the morning sky but not well placed for observation, rising barely an hour before the sun. It is in close conjunction with Mercury on the 10th and with Venus on the 16th.

Uranus on the 15th is in R.A. 3h 3m, Decl. 16° 56' N. and transits at 11.33. Neptune on the 15th is in R.A. 11h 27m, Decl. 4° 48' N. and transits at 19.56. Pluto—For information regarding this planet, see p. 23.
				MAY		Config.
				75th Meridian Civil Time	Min. of Algol	Jupiter's Sat. 4h 45m
	d	h	m		h m	
Mon.	1	4		§ Greatest elongation W., 26° 55'.		
Tue.	2			······································	. 22 59	21043
Wed.	3	10	15	Total Eclipse of (, see p. 00 Full Moon.		32401
Thu.	4	10	10	Eta Aquarid Meteors, p. 54		34102
Fri.	5					d4301
Sat.	6					42310
Sun.	7					d4O23
Mon.	8					40123
Tue.	9	4		₫δ⊙		42103
I uc.	0			♂♂℃ ♂ 6° 53′ S.	•	12100
Wed.	10	10		σ ^φ ^μ ^β ^μ ^β		42301
Thu.	11	0		Moon in Apogee. Dist. from \oplus , 251,200 mi		31402
i nu.	11	-	40		.10 20	01402
Fri.	12	0	10			30214
Sat.	13					23104
Sun.		15		Greatest Hel. Lat. S. Greatest H		01234
Sun.	17		03	$\sigma' 2 \mathbb{Q}$ 24 4° 36′ S.	. 10 15	01204
Mon.	15	10	00			0234*
Tue.		16		σ ♀ þ ♀ 0° 34′ N		21034
i uc.	10		15	$\sigma \flat \mathbb{C}$ \flat 3° 36' S.		21001
				$\sigma' \neq \mathbb{Q} \qquad \qquad \varphi \qquad $		
Wed.	17			σξ (ξ 2° 54′ S	7 04	d2O14
Thu.	18			ở δ ⊈ 6 0° 36′ N		31024
I nu.	. 10		25			01021
Fri.	19	4	-0	Q Greatest Hel. Lat. S.		30241
Sat.	20	-				23140
Sun.	21					4013*
Mon.						41023
Tue.	$\overline{23}$	7		Moon in Perigee. Dist. from⊕, 228,800 mi		42103
Wed.				······································		42031
Thu.	25	3		σ 및 Ô 및 1° 12′ S		43102
		-	20			
Fri.	26					43021
Sat.	27	2	27	σΨ @ Ψ 5° 21′ N		43210
Sun.	28					031**
Mon.	29					10243
Tue.	30					d2O34
Wed.	31				15 09	20134

THE SKY FOR JUNE, 1939

The times of transit are given in local mean time, 0h at midnight; to change to Standard Time, see p. 10. Estimates of altitude are for an observer in latitude 45° N.

The Sun—During June the sun's R.A. increases from 4h 32m to 6h 36m and its Decl. changes from 21° 54' N. to a maximum value of 23° 26' 40" N. on the 22nd and then drops back to 23° 11' N. at the end of the month. The equation of time drops from +2m 30s to -3m 27s (see p 7). The sun is at the summer solstice, the point in its path farthest north of the equator, at 7h 40m G.C.T., June 22. At this time it enters the sign Cancer and summer commences. The duration of daylight is now greatest and does not change appreciably for some days, see p. 13. During the last part of June and first part of July the local mean time of sunset appears almost constant owing to the decrease in the equation of time.

The Moon—For its phases, perigee and apogee times and distances, and its conjunctions with the planets, see opposite page.

Mercury on the 15th is in R.A. 6h 10m, Decl. $25^{\circ}8'$ N. and transits at 12.44. It is in superior conjunction with the sun on the 7th and not well placed for observation during June.

Venus on the 15th is in R.A. 3h 55m, Decl. 19° 9' N. and transits at 10.26. It is a bright yellow star in the morning sky but is becoming less favourably placed for observation as it approaches superior conjunction and rises barely an hour before sunrise.

Mars on the 15th is in R.A. 20h 31m, Decl. $22^{\circ} 54'$ S. and transits at 3.00. It is a red star, becoming noticeably brighter as it moves into the evening sky. It is now of magnitude -1.5, being of the same brightness as Sirius. On June 24th it reaches a stationary point in its path among the stars and starts to retrograde, or move westward.

Jupiter on the 15th is in R.A. 0h 23m, Decl. 1° 8' N. and transits at 6.52. It is in quadrature on June 29 and is visible in the morning sky during the last half of the night. Its magnitude is now -2. For the configurations of its satellites, see opposite page, and for their eclipses, etc., see p. 50.

Saturn on the 15th is in R.A. 1h 49m, Decl. $8^{\circ} 42'$ N. and transits at 8.17. It is slowly separating from the sun in the morning sky, rising just under 3 hours before the sun. It is a pale yellow star of magnitude 0.7. For the elongations, etc., of its satellites, see p. 52.

Uranus on the 15th is in R.A. 3h 10m, Decl. 17° 24' N. and transits at 9.38. Neptune on the 15th is in R.A. 11h 27m, Decl. 4° 48' N. and transits at 17.54. Pluto—For information regarding this planet, see p. 23.

			JUNE			Config.
				Mir		Jupiter's
			75th Meridian Civil Time	of Alg		Sat. 3h 30m
		n m		h	m	
Thu.		2 11	•			31024
Fri.	2	-	Ψ Stationary in R.A			30124
		6	ų in Ω			
Sat.	3		•••••••••••••••••••••••••••••••••••••••		58	32104
Sun.	4					2014*
Mon.		4	σΥδ ♀ 1°11′S			10423
Tue.			రరి⊄ రి 8°39′S		46	d4O13
Wed.	-	4				42O3*
		7	§ in Perihelion			
		8	Moon in Apogee. Dist. from \oplus , 251,400 mi.			
Thu.	8		·····			43102
Fri.		3 07			35	43012
Sat.	10 1	-	b Greatest Hel. Lat. S			43210
Sun.			o 24 € 24 4° 12′ S			4201*
Mon.		6	□Ψ⊙		24	41023
Tue.			σ 𝑘 𝔅 🖞 🖞 δ			40213
Wed.			් ර ් ී € ර ් ී € 0° 47′ N		13	21043
Thu.		6 16	ଦ′ହ ℂ ହ 0° 58′ N			d3O24
Fri.	16		•••••••••••••••••••••••••••••••••••••••			30124
Sat.		8 37	-	.20	02	32104
	-	3				
Sun.			୦ ୪୪ ଏ ଓ ଓ 6° 28′ N			23014
Mon.	19 1	5	Moon in Perigee. Dist. from \oplus , 225,800 mi			10234
Tue.	20		••••••••••••••••••		50	O2134
Wed.						21034
Thu.			\bigcirc enters \textcircled{O} , Summer commences. Long. of \bigcirc , 90			d 0 14*
Fri.				13 3	39	34O2*
9		3 35	~			
Sat.		3	o ⁷ Stationary in R.A			43210
Sun.	25		•••••••••••••••••••••••••••••••••••••••			42301
Mon.			••••••		28	41032
Tue.	27		·····			40213
Wed.	28	•				42103
Thu.	29 2	U	$\Box 2 \odot \dots$	•••	16	4031*
Fri.	30					302**

The times of transit are given in local mean time, 0h at midnight; to convert to Standard Time, see p. 10. Estimates of altitude are for an observer in latitude 45° N.

The Sun—During July the sun's R.A. increases from 6h 36m to 8h 41m and its Decl. changes from 23° 11' N. to 18° 19' N. The equation of time drops from -3m 27s to a minimum of -6m 24s on the 27th and then rises to -6m 16s at the end of the month (see p. 7). For changes in the length of the day, see p. 14. The sun enters Leo, the second summer sign of the zodiac, on July 23rd. The earth is in aphelion, the point in its orbit farthest from the sun, on the 5th of the month. For distance from the sun, see opposite page.

The Moon—For its phases, perigee and apogee times and distances, and its conjunctions with the planets, see opposite page.

Mercury on the 15th is in R.A. 9h 21m, Decl. $14^{\circ} 42'$ N. and transits at 13.53. It reaches greatest elongation east of the sun on the 13th and is fairly well placed for observation in the evening sky, setting over an hour after the sun. It is a star of magnitude 0.7, 13° above the horizon at sunset.

Venus on the 15th is in R.A. 6h 31m, Decl. 23° 16' N. and transits at 11.05. It is in the morning sky, rising shortly before the sun and not particularly well placed for observation. It is of stellar magnitude -3.3.

Mars on the 15th is in R.A. 20h 22m, Decl. 25° 33' S. and transits at 0.53. It is in opposition to the sun on July 23 and is in view all night at this time. It has now reached a stellar magnitude -2.6, which is as bright as Jupiter ever appears. This opposition and the equally favourable one in 1941 will be the two best oppositions for Martian observation in the next 15 years. At such times the brightness of the planet is rather surprising to the layman. Mars is closest to the earth a few days after opposition, on July 27. It is 36,000,000 miles distant at this time and has an apparent angular diameter of 24 seconds of arc.

Jupiter on the 15th is in R.A. 0h 33m, Decl. 2° 5' N. and transits at 5.04. It rises in the east shortly before midnight and is a brilliant star visible for the last half of the night. On July 30 it is at a stationary point in its path among the stars and starts to move westward or retrograde. For the configurations of its satellites, see opposite page, and for their eclipses, etc., see p. 50.

Saturn on the 15th is in R.A. 1h 57m, Decl. $9^{\circ} 21'$ N. and transits at 6.27. It is in the morning sky, rising about 5 hours before the sun. Quadrature is reached on the 24th. For the elongations, etc., of its satellite, see p. 52.

Uranus on the 15th is in R.A. 3h 15m, Decl. 17° 45' N. and transits at 7.46. Neptune on the 15th is in R.A. 11h 29m, Decl. 4° 37' N. and transits at 15.58. Pluto—For information regarding this planet, see p. 23.

				JULY			Config.
						in.	Jupiter's
				75th Meridian Civil Time		f gol	Sat. 2h 45h
	d	h	m		h	m	
Sat.	1	11	16	Full Moon			dd3O4
Sun.	2					05	23014
Mon.	3	17	56	ଟ ଟି ଏ ସି 10° 36′ S			10324
Tue.	4			· · · · · · · · · · · · · · · · · · ·			01234
Wed.	5	9		Moon in Apogee. Dist. from⊕, 252,000 mi	. 0	54	21034
		15		\oplus in Aphelion. Dist from \bigcirc , 94,452,000 mi.			
Thu.	6						20134
Fri.	7				. 21	42	31024
Sat.	8						30214
Sun.	9	0	58	σ 24 € 24 3° 52′ S			32041
		14	49	C Last Quarter.			
Mon.	10	18	58	♂ 𝑘 𝔅 ¬ 𝑘 3° 00′ S	. 18	31	41032
Tue.	11	0		₿ in ♡			40123
Wed.	12	6	57	ර ී € ී 1° 03′ N			412O3
Thu.	13	14		Greatest elongation E. 26° 31'	. 15	20	42013
Fri.		11		♀ in &			43102
Sat.	15	16	14	ଏହି ⊈ ଦୁ 4° 33′ N			43012
Sun.	16	16	03	New Moon	. 12	08	4320*
Mon.	17	18		Moon in Perigee. Dist. from \oplus , 223,200 mi			4130*
Tue.	18	13	35	∀ [₿] ^{3°} 25′ N			01423
Wed.	19			•••••••••••••••••••••••••••••••••••••••	. 8	57	12034
Thu.	20	15	40	∀Ψ 4° 53′ N			20134
Fri.	21	6		§ in Aphelion			13024
Sat.	22				. 5	45	30124
Sun.	23	3		$\circ^{\circ} \circ^{\uparrow} \odot$ Dist. from \oplus , 36,170,000 mi			3204*
		6	34				
Mon.	24	13		□ ▶ ⊙			d 04**
Tue.	25			· · · · · · · · · · · · · · · ·		34	01423
Wed.	26	17		§ Stationary in R.A			12403
Thu.	27	16				23	42013
Fri.	28			Delta Aquarid Meteors, p. 54			41302
Sat.	29						43012
Sun.	30	7		24 Stationary in R.A		11	43210
/		11	13	ଟ ଟି ⊈ ଟି 11° 41′ S.	-		
Mon.	31		37				43201

The times of transit are given in local mean time, 0h at midnight; to change to Standard Time, see p. 10. Estimates of altitude are for an observer in latitude 45° N.

The Sun—During August the sun's R.A. increases from 8h 41m to 10h 37m and its Decl. changes from $18^{\circ} 19'$ N. to $8^{\circ} 42'$ N. The equation of time increases from -6m 16s to -0m 20s (see p. 7). For changes in the length of the day, see p. 14. The sun enters Virgo, the third summer zodiacal sign, on the 23rd.

The Moon—For its phases, perigee and apogee times and distances, and its conjunctions with the planets, see opposite page.

Mercury on the 15th is in R.A. 9h 1m, Decl. $12^{\circ} 43'$ N. and transits at 11.27. It is in inferior conjunction with the sun on the 10th. August 28, Mercury reaches greatest elongation west of the sun in the morning sky, at which time it rises in the east about an hour and a half before the sun. It will be visible for some days before and after the 28th as a reddish star of 0 magnitude.

Venus on the 15th is in R.A. 9h 12m, Decl. 17° 18' N. and transits at 11.43. It is approaching the sun in the morning sky and not well placed for observation.

Mars on the 15th is in R.A. 19h 51m, Decl. $27^{\circ} 23'$ S. and transits at 22.15. It is a very prominent red star in the evening sky, appearing in the south-east directly after sunset. On August 24 it reaches a stationary point in its path as it ceases to retrograde, commencing to move eastward once more among the stars.

Jupiter on the 15th is in R.A. 0h 33m, Decl. 1° 56' N. and transits at 3.02. The planet rises in the east just under two hours after sunset and during the last month has materially increased in brightness, being now of magnitude -2.4. At this time both Jupiter and Mars are brilliant objects in the evening and add considerably to the beauty of the summer sky, making it this year the rival of the brilliant winter groups of stars. For the configurations of Jupiter's satellites see opposite page, and for their eclipses, etc., see p. 50.

Saturn on the 15th is in R.A. 2h 0m, Decl. 9° 29' N. and transits at 4.29. It is moving into the evening sky, rising shortly before midnight. It has increased in brightness to magnitude 0.4 as opposition is approached. A stationary point in its path among the stars is reached on August 14, on which date it commences to retrograde. For the elongations of its satellites, etc., see p. 52.

Uranus on the 15th is in R.A. 3h 18m, Decl. 17° 56' N. and transits at 5.46. *Neptune* on the 15th is in R.A. 11h 32m, Decl. 4° 16' N. and transits at 13.59. *Pluto*—For information in regard to this planet, see p. 23.

				AUGUST	n.	Config. of Jupiter's
				75th Meridian Civil Time Alg	f	Sat. 2h 00m
	d	h	m	h	m	
Tue.	1	19		Moon in Apogee. Dist. from \oplus , 252,400 mi		4032*
Wed.	2				00	d4103
Thu.	3					24013
Fri.	4					d1024
Sat.	5	8	46		48	30124
Sun.	6					32104
Mon.	7					32014
Tue.	8	4	18	Last Quarter	37	O324*
		16	28	♂ ô € ô 1° 21′ N.		
Wed.	9			· · · · · · · · · · · · · · · · · · ·		10234
Thu.	10	11		$\sigma \notin \odot$ Inferior		20134
		14				
Fri.	11			Perseid Meteors, p. 54 7	25	10324
Sat.	12					34012
Sun.	13	6		280 8 E ^o 40/ S		43210
Mon.	14	11	52	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	14	43201
		15	54	o		
		18		b Stationary in R.A.		
		22	53	•		
Tue.	15	3		Moon in Perigee. Dist. from \oplus , 222,000 mi		41032
		10				
Wed.	16			·		d4O23
Thu.	17	2	03	$\sigma' \Psi \mathbb{G} \qquad \qquad \Psi \qquad 4^{\circ} 38' \mathrm{N} 1$	03	42013
		3		♀ in Perihelion.		
Fri.	18			•,••••••••••••••		41023
Sat.	19	20		§ Stationary in R.A21	51	34012
Sun.	20					31204
Mon.		16	21	First Quarter		32014
Tue.	22				40	10324
	23			o^{7} Greatest Hel. Lat. S		01234
Thu.	24	10		o ⁷ Stationary in R.A		2034*
Fri.	25				28	1034*
Sat.	26	9	02	o′ o [¬] € o [¬] 11° 00′ S		30124
Sun.	27					31204
Mon.	28	2		§Greatest elongation W., 18° 16'12	17	32041
		11		Stationary in R.A.		
		22		Moon in Apogee. Dist. from \oplus , 252,500 mi.		
Tue.	29	16		ξ ² in Ω		4102*
		17	09	Full Moon.		
						40123
Wed.	30					40123

THE SKY FOR SEPTEMBER, 1939

The times of transit are given in local mean time, 0h at midnight; to change to Standard Time, see p. 10. Estimates of altitude are for an observer in latitude 45° N.

The Sun—During September the sun's R.A. increases from 10h 37m to 12h 25m and its Decl. changes from $8^{\circ} 42'$ N. to $2^{\circ} 45'$ S. The equation of time increases from -0m 20s to +9m 55s (see p. 7). On the 23rd, at 22h 50m G.C.T., the sun is at the autumnal equinox and crosses the equator on its way south, entering the sign Libra. Autumn commences at this time and days and nights are approximately equal all over the world.

The Moon—For its phases, perigee and apogee times and distances, and its conjunctions with the planets, see opposite page.

Mercury on the 15th is in R.A. 11h 6m, Decl. 7° 46' N. and transits at 11.36. It is visible in the morning sky just before sunrise for a few days at the beginning of the month. It then slips back into the twilight and is in superior conjunction with the sun on the 22nd.

Venus on the 15th is in R.A. 11h 39m, Decl. $3^{\circ} 46'$ N. and transits at 12.07. It is in superior conjunction with the sun on the 5th, at which time it passes into the evening sky. It is too near the sun for observation in September.

Mars on the 15th is in R.A. 20h 01m, Decl. $25^{\circ} 21'$ S. and transits at 20.25. It is slowly fading as the distance between it and the earth increases. It is still very prominent in the evening sky, however, being a red star of magnitude -1.5.

Jupiter on the 15th is in R.A. 0h 23m, Decl. 0° 41' N. and transits at 0.49. It reaches opposition on the 27th and is in view all night throughout the month. This is the most favourable opposition of Jupiter for the next 12 years since the planet is nearest its perihelion point at this opposition. Because of this fact it is considerably nearer the earth than at the usual opposition, being 367,000,000 miles distant. This is 47,000,000 miles closer than it was at the opposition of 1934, for example. Jupiter reaches its maximum brightness this month at magnitude -2.5. For the configurations of its satellites, see opposite page, and for their eclipses, etc., see p. 50.

Saturn on the 15th is in R.A. 1h 57m, Decl. $9^{\circ} 4'$ N. and transits at 2.24. It is approaching opposition to the sun and is in view most of the night. For the elongations, etc., of its satellites, see p. 52.

Uranus on the 15th is in R.A. 3h 18m, Decl. 17° 55' N. and transits at 3.44. *Neptune* on the 15th is in R.A. 11h 36m, Decl. 3° 50' N. and transits at 12.01. *Pluto*—For information regarding this planet, see p. 23.

				SEPTEMBER	Min.	Config. of Jupiter's
				75th Meridian Civil Time	of Algol	Sat. 1h 15m
	d	h	m		h m	
Fri.	1	11	14	o 24 € 24 3° 48′ S		42103
Sat.	2			· · · · · · · · · · · · · · · · · · ·		43012
Sun.	3	6		§ in Perihelion		43120
		9	23	♂ þ ℂ þ 2° 42′ S.		
Mon.	4			άδα δ 1° 36′ Ν		43201
Tue.		15		$\sigma \neq \odot$ Superior		14302
Wed.	6	15	24			01423
Thu.		23		9 Greatest Hel. Lat. N		21043
Fri.	8				23 31	d2O34
Sat.	9					30124
Sun.	10					d3104
Mon.	11				20 20	32014
Tue.	12	13		Moon in Perigree. Dist. from \oplus , 222,500 mi		13024
			47	σ'⊈		
Wed.	13		22			01423
		13		g Greatest Hel. Lat. N.		
		13	13	σ′♀ € ♀ 4° 49′ N.		
		14	15	σΨ C Ψ 4° 30′ N.		
Thu.	14			$\sigma \varphi \Psi \qquad \varphi \qquad 0^{\circ} 16' \text{ N.} \dots$	17 08	214O3
Fri.	15			•••••		42013
Sat.	16	16		σΨΟ		4302*
Sun.	17	13		σ^1 in Perihelion		43102
Mon.	18			-		43201
Tue.	19	7		σ ⁽		41302
Wed.	20		34		10 46	40132
Thu.	21			~		42103
Fri.	22	9		σ 𝔅⊙ Superior		20413
Sat.	23	3	05	♂♂℃ ♂ 9° 42′ S	7 34	dO24*
		17	50	\odot enters \simeq , Autumn comm. Long. of \odot 180	•	
Sun.	24			· · · · · · · · · · · · · · · · · · ·		31024
Mon.	25	4		Moon in Apogee. Dist from ⊕, 252,200 mi		32014
Tue.	26	-		······································		3104*
Wed.		14		$^{\circ}2$ Dist. from \oplus , 367,100,000 mi.		O3124
Thu.	28		27			12034
	-•			σ21 € 24 4° 06′ S.		
Fri.	29	- •		·····	1 11	20143
Sat.		12	19	♂▶ € þ 2° 50′ S		10342

THE SKY FOR OCTOBER, 1939

The times of transit are given in local mean time, 0h at midnight; to change to Standard Time, see p. 10. Estimates of altitude are for an observer in latitude 45° N.

The Sun—During October the sun's R.A. increases from 12h 25m to 14h 21m and its Decl. changes from 2° 45' S. to 14° 4' S. The equation of time rises from +9m 55s to +16m 20s (see p. 7). For changes in the length of the day see p. 15. The sun enters Scorpio, the second autumnal sign of the zodaic, on the 24th of the month. There is a total eclipse of the sun on October 12, invisible in North America. For details, see p. 25.

The Moon—For its phases, perigee and apogee times and distances, and its conjunctions with the planets, see opposite page. There is a partial eclipse of the moon on October 28, visible in Canada and the United States. For details see p. 25.

Mercury on the 15th is in R.A. 14h 13m, Decl. 14° 19' S. and transits at 12.43. It is too near the sun for observation during October.

Venus on the 15th is in R.A. 13h 57m, Decl. 11° 10′ S. and transits at 12.27. It is in the evening sky but sets only a few minutes after the sun and so is not well placed for observation.

Mars on the 15th is in R.A. 20h 50m, Decl. 20° 54' S. and transits at 19.18. It is a red star of magnitude -0.7 in the evening sky, setting over six hours after the sun.

Jupiter on the 15th is in R.A. 0h 8m, Decl. 0° 52' S. and transits at 22.33. It is a brilliant star of magnitude -2.4, appearing in the east at sunset and remaining in view all night. For the configurations of its satellites, see next page, and for their eclipses, etc., see p. 50.

Saturn on the 15th is in R.A. 1h 49m, Decl. 8° 18' N. and transits at 0.18. It is in opposition to the sun on the 22nd and is in view all night throughout the month. It reaches a maximum brightness at this time of magnitude 0.1. The rings of Saturn at this opposition are opened a little more than half the maximum possible value. They appear inclined to the line of sight by about 15 degrees. For some years now the brightness of the planet has increased at each opposition owing to the opening out of the rings as seen from the earth. These will appear at their maximum inclination to the line of sight at the end of 1943. For the elongations, etc., of the satellites of Saturn, see p. 52.

Uranus on the 15th is in R.A. 3h 15m, Decl. 17° 43' N. and transits at 1.43. Neptune on the 15th is in R.A. 11h 40m, Decl. 3° 25' N. and transits at 10.07. Pluto—For information regarding this planet, see p. 23.

				OCTOBER		Config.
					Min.	Jupiter's
				75th Meridian Civil Time	of Algol	Sat. 23h 30m
	d	h	m		h m	
Sun.	1	20		σ [′] ^β ♀ ^β 0° 36′ S		4320*
Mon.	2	3	49	♂ ô € ô 1° 41′ N		43120
Tue.	3					40312
Wed.	4				18 49	412O3
Thu.	5					42013
Fri.	6	0	27	C Last Quarter		41032
		23		₿ in °°		
Sat.	- 7				15 38	34012
Sun.	8					3204*
Mon.	9					32104
Tue.	10	20		Moon in Perigee. Dist. from \oplus , 224,700 mi		0124*
Wed.	11	2	21	$\sigma' \Psi \mathbb{G} \qquad \Psi \qquad 4^{\circ} 27' \mathrm{N}. \dots \dots \dots$		d1034
Thu.	12			Total Eclipse of \bigcirc , see p. 25		20134
		15	30			
Fri.	13	9	45		9 15	10234
		16	27	σ ['] ξ C ^β 1° 17′ S.		
Sat.	14					30124
Sun.	15					32104
Mon.	16				6 04	d324O
Tue.	17	5		§ in Aphelion		43012
Wed.	18					41023
Thu.	19	22	24	First Quarter. Orionid Meteors, p. 54	252	42013
Fri.	20					41023
Sat.	21	13	09	$\sigma' \sigma^{?} \textcircled{6}$ $\sigma' \ 8^{\circ} 11' \text{ S}$ $\rho \flat \odot$ Dist. from \oplus , 771,900,000 mi.	$23 \ 41$	43012
		22		$\circ^{\circ} \mathfrak{b} \odot$ Dist. from \oplus , 771,900,000 mi.		
Sun.	22	18		Moon in Apogee. Dist. from \oplus , 251,600 mi		43210
Mon.	23					34201
Tue.	24				$20 \ 30$	3402*
Wed.	25	11	42	of 24 € 24 4° 22′ S		10243
Thu.	26					20134
Fri.	27	15	09	♂ 𝑘 𝔅 🕴 3° 03′ S	$17 \ 19$	1034*
Sat.	28			Partial Eclipse of C, see p. 25		30124
		- 1	42			
Sun.	29	8	13	♂δ € δ 1° 37′ N		31204
Mon.	30			· · · · · · · · · · · · · · · · · · ·	$14\ 07$	32014
Tue.	31					3024*

The times of transit are given in local mean time, 0h at midnight; to change to Standard Time, see p. 10. Estimates of altitude are for an observer in latitude 45° N.

The Sun—During November the sun's R.A. increases from 14h 21m to 16h 24m and its Decl. changes from $14^{\circ}4'$ S. to $21^{\circ}38'$ S. The equation of time increases from +16m 20s to a maximum of +16m 23s on November 4 and then decreases to +11m 22s at the end of the month (see p. 7). For changes in the length of the day, see p. 16. On the 23rd the sun enters Sagittarius, the third autumn sign of the zodiac.

The Moon—For its phases, perigee and apogee times and distances, and its conjunctions with the planets, see opposite page.

Mercury on the 15th is in R.A. 16h 45m, Decl. $24^{\circ} 49'$ S. and transits at 13.11. On November 7 it is at greatest elongation east of the sun in the evening sky but is not well placed for observation, setting barely 30 minutes after the sun. It is in inferior conjunction on the 28th.

Venus on the 15th is in R.A. 16h 32m, Decl. $22^{\circ} 25'$ S. and transits at 13.01. It sets about an hour after the sun in the evening sky.

Mars on the 15th is in R.A. 22h 0m, Decl. 14° 7' S. and transits at 18.25. It is in quadrature on November 29 and appears in the western evening sky for the first half of the night. It is now of zero magnitude.

Jupiter on the 15th is in R.A. 23h 59m, Decl. 1° 46' S. and transits at 20.22. It has faded slightly as it recedes from the earth after opposition but is still a very brilliant object in view for most of the night. It is at a stationary point on the 25th and commences to move eastward among the stars again on this date. For the configurations of its satellites, see opposite page, and for their eclipses, etc., see p. 50.

Saturn on the 15th is in R.A. 1h 40m, Decl. 7° 29' N. and transits at 22.03. It is well placed for observation, rising in the east shortly before sunset and being visible most of the night. For the elongations, etc., of its satellites, see p. 52.

Uranus on the 15th is in R.A. 3h 10m, Decl. 17° 24' N. and transits at 23.32. Opposition to the sun occurs on the 13th.

Neptune on the 15th is in R.A. 11h 43m, Decl. 3° 4' N. and transits at 8.09. *Pluro*—For information regarding this planet, see p. 23.

				NOVEMBER	Config.
				Min.	Jupiter's
				75th Meridian Civil Time of Algol	Sat. 22h 00m
	d	h	m	h m	
Wed.	1				dO243
Thu.	2				24013
Fri.	3	1		ዩ in የየ	41203
Sat.	4	8	12	Last Quarter	d4012
Sun.	5				4312O
Mon.	6	13		g Greatest Hel. Lat. S	43201
Tue.	7	12	18	$ \vec{\nabla} \Psi $ Ψ $4^{\circ} 23' N.$	43102
		16		Moon in Perigee. Dist. from \oplus , 228,100 mi	
		22		§ Greatest elongation E., 23° 10'.	
Wed.	8				d4O32
Thu.	9				4203*
Fri.	10				21043
Sat.	11		54	W New Moon 1 23	O3124
Sun.	12	10	42	σ ♀ ① ♀ 3° 37′ S	d3104
Sum				σ ⊈ ⊈ 6° 12′ S.	
Mon.	13	1			32014
Tue.	14				31024
Wed.				Leonid Meteors, p. 54	01324
Thu.	16				2034*
Fri.	17			·····	21043
Sat.	18	1		σ'⋭♀ ₿ 1°24' S	04312
5411		9		Stationary in R.A.	
		18	21	First Quarter.	
Sun.	19	7	22	$\sigma' \sigma' $ $\sigma' = 6^{\circ} 23' S 15 50$	34102
		14		Moon in Apogee. Dist. from \oplus , 251,200 mi.	
Mon.	20				43201
Tue.		16	49	σ´2↓ € 2↓ 4° 20′ S	4310*
Wed.					4012*
Thu.	23	19	49	♂ 𝔥 𝔄 👂 3° 09′ S	42103
Fri.	24				d42O3
Sat.	$\overline{25}$	6		24 Stationary in R.A 9 28	40132
Sur.		14		ğ in Ω	
			07	σ΄δ ⊈ δ 1° 31′ N.	
Sun.	26		54		34102
Mon.					32041
Tue.	28	12		$\sigma \notin \bigcirc$ Inferior	3104*
Wed.	29	3		$\Box \sigma^{2} \Theta \dots $	0124*
Thu.	30	5		۵ in Perihelion	21034
	00			+ In termenon	

THE SKY FOR DECEMBER, 1939

The times of transit are given in local mean time, 0h at midnight; to change to Standard Time, see p. 10. Estimates of altitude are for an observer in latitude 45° N.

The Sun—During December the sun's R.A. increases from 16h 24m to 18h 41m and its Decl. changes from 21° 38'S. to a minimum value of 23° 26' 40"S. on the 22nd, rising to 23° 7'S. at the end of the month. The equation of time drops from +11m 22s to -2m 58s (see p. 7). At 18h 6m G.C.T., December 22, the sun enters Capricornus, the first winter sign of the zodiac, and winter commences. It is at its position farthest south of the equator and the length of daylight in the northern hemisphere is at a minimum, changing very slightly for several days, see p. 16.

The Moon—For its phases, perigee and apogee times and distances, and its conjunctions with the planets, see opposite page.

Mercury on the 15th is in R.A. 15h 57m, Decl. 17° 59' S. and transits at 10.26. It reaches greatest elongation west of the sun in the morning sky on December 16, rising an hour and three-quarters before the sun at this time. It will be in the S.E. and 16° above the horizon at sunrise.

Venus on the 15th is in R.A. 19h 15m, Decl. $23^{\circ} 51'$ S. and transits at 13.46. It is a bright yellow star in the evening sky, setting in the south-west almost two hours after the sun.

Mars on the 15th is in R.A. 23h 12m, Decl. $6^{\circ} 2'$ S. and transits at 17.39. It is moving north among the stars of the evening sky and sets about seven hours after the sun, being in view for the first half of the night. It is growing fainter as the distance between Mars and the earth increases.

Jupiter on the 15th is in R.A. 0h 1m, Decl. 1° 26' S. and transits at 18.26. It is in the south-east at sunset and sets shortly after midnight. Quadrature with the sun is reached on the 22nd. For the configurations of its satellites, see opposite page, and for their eclipses, etc., see p. 50.

Saturn on the 15th is in R.A. 1h 35m, Decl. 7° 4' N. and transits at 20.00. It is in view most of the night, appearing in the south-east after sunset. Its magnitude has now dropped to 0.5 as it recedes from the earth after opposition. A stationary point in its path among the stars is reached on the 28th, at which time it resumes its eastward motion.

Uranus on the 15th is in R.A. 3h 5m, Decl. 17° 5' N. and transits at 21.30. Neptune on the 15th is in R.A. 11h 45m, Decl. 2° 54' N. and transits at 6.12. Pluto—For information regarding this planet, see p. 23.

				DECEMBER		Config.
				75th Meridian Civil Time	Min. of Algol	Jupiter's Sat. 20h 45m
	d	h	m		h m	
Fri.	1				. 3 06	20134
Sat.	2				•	O234*
Sun.	3	0		24 Greatest Hel. Lat. S	.23 55	31024
		$\frac{2}{15}$	20	Moon in Perigee. Dist. from⊕, 230,100 mi. ℂ Last Quarter.		
Mon.	4	19	40	$\sigma \Psi \mathbb{Q} \Psi 4^{\circ} 14' \text{ N}$		32014
Tue.	5					31204
Wed.	6				.20 44	43012
Thu.	7	10		Q in Alphelion		412O3
Fri.	8	1		§ Stationary in R.A		42013
Sat.	9	4	47	σ ['] ^β [©] ^β ^{0°} 13' N		41023
Sun.	10	12		§ Greatest Hel. Lat. N		43102
		16	45	New Moon.		
Mon.	11			 		43201
Tue.	12	19	48			43120
Wed.	13					34012
Thu.	14					d1043
Fri.	15					20143
Sat.		19		Greatest elongation W., 21° 25′		10234
Sun.		11		Moon in Apogee. Dist. from \oplus , 251,300 mi		dd O2 4
Mon.		3		$\Box \Psi \odot \dots \dots \square \Psi \odot \dots \dots \dots \square \dots \square \dots \square \dots \square \dots \square \dots \dots \square \dots \dots \square \dots \dots$		3204*
wion.	10	~		ୁ କୁ ପ୍ରୁ ସୁ ସୁ ସୁ 4° 05′ S.	. 0 00	0401
			04			
Tue.	19			$\sigma' 2 \mathbb{Q}$ 24 3° 55′ S		32104
Wed.	20	2	00			30124
Thu.	20	2	52	♂ 𝔥 𝔄 𝔥 2° 59′ S	. 4 49	10234
Fri.				\odot enters \eth , Winter comm. Long. of \odot , 270°.		24013
1.11.	22	18		$\Box 2 \odot$	•	21010
				$\sigma \circ \mathbb{C}$ \circ 1° 32′ N.		
Sat.	23	21	10			4103*
Sun.	23 24					40312
Mon.						4320*
Tue.	23 26	в	28	Full Moon		43210
Wed.		U	20			43012
Thu.		21		b Stationary in R.A		4102*
Fri.	29	0		Ψ Stationary in R.A		42013
	20	6		Moon in Perigee. Dist. from \oplus , 227,300 mi.		12010
		21		$\label{eq:point} \end{tabular}$		
				• • • • • • • • • • • • • • • • • • • •		1400*
Sat.	30					1403*

PHENOMENA OF JUPITER'S SATELLITES, 1939

E-eclipse, O-occultation, T-transit, S-shadow, D-disappearance, R-reappearance, I-ingress, e-egress. The Roman numerals denote the satellites. 75th Meridian Civil Time. (For other times see p. 00).

		1								
JANUARY			•			JULY	-Ca	on.		
d h m Sat. Phen. d h m Sat. Ph	hen. ER	d	h	m	Sat.		d	h m		Phen.
4 18 45 IV TI 19 44 I 5 20 05 I OD 16 18 05 II	ER	30 31	04 01	$\frac{25}{39}$	I I	ED SI		$\begin{array}{ccc} 02 & 52 \\ 03 & 52 \end{array}$	I	TI Se
6 18 20 I SI 21 18 38 I	OD	91	$01 \\ 02$	39 45	İI	ED		$ \begin{array}{ccccccccccccccccccccccccccccccccccc$	İ	ED
19 39 I Te 22 18 13 I	Te	=								
7 17 47 II SI 18 57 I 17 48 I ER 25 18 07 III	Se OD					AUG	GUS	г		
18 43 II Te 29 17 58 I	ΤI	d	h	m		Phen.	d	h m		Phen.
18 52 III TI 18 36 I 13 19 24 I TI 30 18 02 I	SI ER	1	$\frac{02}{23}$	$\frac{20}{30}$	I I	OR Te	17	$\begin{array}{ccc} 00 & 23 \\ 21 & 32 \end{array}$	I	OR Te
13 19 24 I TI 30 18 02 I 14 18 42 II TI	ĿК		$\frac{23}{28}$	59	ÎΙ	ŤĬ	18	$01 \ 42$	ĪI	OR
		2	00	$12_{0.5}$	ΙI	Se	01	$\begin{array}{ccc} 03 & 24 \\ 21 & 22 \end{array}$	III III	SI OD
FEBRUARY		4		$\frac{35}{22}$	II III	Te TI	21	23 46	III	OR
d h m Sat. Phen.		-	$ \begin{array}{c} 02 \\ 03 \end{array} $	$\frac{47}{33}$	III I	Te SI		$\begin{array}{c} 04 & 37 \\ 01 & 49 \end{array}$	I	ED SI
5 18 13 III Se		1 1	03	33 40	İ	TI	20	$01 49 \\ 02 40$	I	ΤÎ
Jupiter being near the Sun, phenomena	a of	8	00	48	Ī	ED		04 02	Į	Se
the Satellites are not given from Februar			$\frac{04}{22}$	09 01	I I	OR SI		$\begin{array}{ccc} 04 & 51 \\ 23 & 05 \end{array}$	I I	Te ED
to May 1.			23	07	I	ΤI	24	02 09	I	OR
		9	00	07	ΙI	SI		$ \begin{array}{cccc} 21 & 07 \\ 22 & 30 \end{array} $	I I	TI Se
MAY			$\begin{array}{c} 00\\01 \end{array}$	14 18	I I	Se Te		$\frac{22}{23}$ 18	Î	Te
d h m Sat. Phen. d h m Sat. P 7 03 54 I TI 28 02 55 IV	hen. OD		02	24	II	ΤI		23 46	II	ED
17 03 50 III TI 03 35 II	ED		$\frac{02}{22}$	$\frac{49}{36}$	II I	Se OR	25	$\begin{array}{c} 04 & 01 \\ 21 & 22 \end{array}$	II II	OR Se
22 03 50 I ED 30 03 05 I	SI	10	$\tilde{2}\tilde{3}$	$\overline{22}$	ÎΙ	ÖR		22 54	II	Te
23 03 24 I Se 03 08 II 24 03 15 III SI 31 03 45 I	Te OR		23	24	III	SI	27	21 08	IV IV	SI Se
		11	$ \begin{array}{c} 02 \\ 02 \end{array} $	$\frac{16}{41}$	III IV	Se SI	28	$\begin{array}{ccc} 22 & 28 \\ 21 & 34 \end{array}$	III	ED
JUNE			03	58	III	ΤI	29	00 26	III	ER
d h m Sat. Phen. d h m Sat. P	hen.	15	$ \begin{array}{c} 04 \\ 02 \end{array} $	$\frac{26}{42}$	IV I	Se ED		$\begin{array}{c} 00 & 47 \\ 03 & 10 \end{array}$		OD OR
4 02 40 III OD 16 02 07 I	OR	10	$2\tilde{3}$	$\frac{1}{55}$	İ	SI	30	$03 \ 42$	I	SI
5 04 06 IV Se 22 01 03 III 6 03 11 II TI 03 14 I	TI SI	16	00	54	I	TI	91	$\begin{array}{c} 04 & 26 \\ 01 & 00 \end{array}$	I	TI ED
03 14 II Se 03 44 III	Ťe		02 02	08 45	I II	Se SI	01	$01 00 \\ 03 55$	Î	OR
7 02 07 I ED 23 04 03 I 8 02 57 I Te 24 01 16 I	OR Te		03	05	I	Te		$22 \ 11$	Ĩ	SI
8 02 57 I Te 24 01 16 I 13 03 08 II SI 29 02 18 III	Se		04	48	II	TI	1	22 52	I	TI
14 04 02 I ED 03 10 II	ED					SEPTI	EME	RER		
15 02 41 I TI 30 02 19 I 03 23 II OR 03 20 IV	ED ED	d	h	m	Sat	Phen.	1 d	h m	Sat.	Phen.
03 34 I Se		ĩ	00	24	I	Se		03 28	II	Te
			$ \begin{array}{c} 01 \\ 02 \end{array} $	$\frac{03}{21}$	I II	Te ED	11	$ \begin{array}{cccccccccccccccccccccccccccccccccccc$	II I	OR ED
JULY			22	$\frac{21}{21}$	Î	ŐŘ	15	01 59	Ī	SI
d h m Sat. Phen. d h m Sat. P		2	21	18	II	SI		02 20	I I	TI
1 00 33 II TI 16 00 36 I 00 59 I TI 04 12 I	ED OR	1	$\frac{22}{23}$	37 59	II II	TI Se		$\begin{array}{c} 04 & 13 \\ 04 & 31 \end{array}$	i	Se Te
01 49 I Se 17 00 00 IV	ΕR	3	01	12	II	Te		19 30	III	SI
03 10 I Te 00 05 I 03 11 II Te 00 24 III	Se ER	5	01 04	$\frac{35}{45}$	III IV	ED ED		$\begin{array}{ccc} 20 & 57 \\ 22 & 17 \end{array}$	III III	TI Se
03 11 II Te 00 24 III 2 00 26 I OR 01 22 I	Te	7	04	40 55	I	ĔD		$\frac{22}{23}$ 18	I	ED
6 03 20 III SI 02 53 II	OR	8	00	05	I	SI		23 19	ĨIJ	Te
7 04 13 I ED 03 05 III 8 00 19 II SI 23 02 30 I	OD ED		00	$\frac{36}{18}$	I I	T I Se	16	$\begin{array}{c} 01 & 50 \\ 20 & 28 \end{array}$	I I	OR SI
01 30 I SI 23 46 I	SI		02	47	I	Te		$20 \ 45$	I	ΤI
02 52 I TI 24 00 10 II	ED		04	56	II	ED		22 41	I	Se
03 01 II Se 01 02 I 03 10 II TI 01 27 III	TI ED		$\frac{20}{21}$	01 23	III I	Te ED	17	$\begin{array}{ccc} 22 & 57 \\ 02 & 33 \end{array}$	I II	Te SI
03 43 I Se 01 58 I	Se	9	00	06	I	OR	- '	03 09	II	ΤI
9 02 19 I OR 03 13 I 10 00 24 II OR 04 24 III	Te ER		$\frac{20}{21}$	$\frac{47}{13}$	I I	Se Te		$\begin{array}{ccc} 05 & 14 \\ 20 & 16 \end{array}$	II I	Se OR
01 48 III OR 25 00 30 I	OR		$\overline{23}$	$\frac{13}{55}$	ÎI	sĩ	18	20 49	II	ED
15 02 56 II SI 26 00 07 II	Te	10	00	54	II	ŤI		23 54	II	OR
03 24 I SI 27 23 08 III	Te	1	02	36	11	Se	1			

	1						
SEPTEMBER—Con.	NOVEMBER—Con.						
d h m Sat. Phen. d h m Sat. Phen. 22 03 54 I SI 05 24 II TI 04 03 I TI 19 42 I ED 23 31 III SI 22 00 I OR	d h m Sat. Phen. d h m Sat. Phen. 22 25 I ER 19 23 52 II TI 10 18 31 I Te 21 22 59 II ER 19 32 I Se 22 19 55 III OR						
23 00 11 III TI 25 19 04 I Se 01 13 I ED 19 06 I Te 02 18 III Se 23 25 II ED	11 23 34 III TI 22 01 III ED 12 21 27 II TI 23 00 38 III ER 23 33 II SI 18 05 II Se						
02 35 III Te 26 02 07 II OR 03 33 I OR 27 18 31 II TI 22 22 I SI 21 06 II Te 22 29 I TI 21 10 II Se	13 00 05 II Te 22 51 I OD 14 20 22 II ER 24 19 59 I TI 15 17 58 III ED 21 11 I SI 20 36 III ER 22 11 I Te						
24 00 36 I Se 30 03 05 I OD 00 40 I Te 03 25 III TI 05 10 II SI 03 33 III SI	23 41 I TI 23 23 I Se 16 00 46 I SI 25 20 44 I ER 21 00 I OD 26 17 52 I Se 17 00 20 I ER 28 20 26 II OD						
OCTOBER	18 08 I TI 29 20 53 III OD						
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	19 15 I SI 23 40 III OR 20 21 I Te 30 18 06 II SI 21 27 I Se 18 13 II Te 18 18 49 I ER 20 42 II Se						
21 31 I OD 19 25 I OD 23 51 I ER 22 10 I ER	DECEMBER						
2 18 38 I TI 23 44 III OD	d h m Sat. Phen. d h m Sat. Phen. 1 00 42 I OD 23 18 II Te						
18 46 I SI 18 18 45 I Te 20 50 I Te 19 17 I Se 20 59 I Se 19 01 16 II TI	21 51 I TI 23 20 II SI 23 07 I SI 16 22 57 I OD						
3 01 45 II OD 02 22 II SI	2 00 03 I Te 17 18 26 III TI 19 10 I OD 20 07 I TI						
18 20 I ER 20 19 22 II OD	22 40 I ER 21 14 III Te 3 17 36 I SI 21 27 I SI						
4 20 45 II TI 21 18 20 III Se	18 28 III Se 22 20 I Te 18 31 I Te 23 39 I Se						
21 07 II SI 22 18 19 II Se 23 21 II Te 23 02 44 I OD	19 47 I Se 18 17 26 I OD 5 22 38 IV TI 19 18 08 I Se						
23 47 II Se 23 51 I TI 8 01 56 I TI 24 00 31 I SI	22 58 II OD 21 23 15 II TI						
02 12 I SI 02 04 I Te 04 08 I Te 02 44 I Se	23 21 IV Te 22 17 24 IV Te 7 18 06 II TI 23 17 26 II OD						
04 25 I Se 21 11 I OD 23 15 I OD 25 00 05 I ER	20 43 II SI 22 51 II ER 20 44 II Te 24 22 03 I TI						
9 01 46 1 ER 03 04 111 OD 20 22 1 TI 18 18 I TI	23 19 II Se 22 28 III TI 8 23 44 I TI 23 24 I SI						
20 41 I SI 19 00 I SI	9 17 34 II ER 25 17 50 II Se 21 03 I OD 19 22 I OD						
22 54 I Se 21 12 I Se	10 18 12 I TI 22 55 I ER 19 32 I SI 26 17 52 I SI						
10 03 59 II OD 26 03 34 II TI 20 15 I ER 18 34 I ER	19 58 III SI 18 44 I Te 20 25 I Te 20 04 I Se						
20 27 III OD 27 21 39 II OD 11 00 30 III ER 28 01 48 II ER	21 43 I Se 27 17 24 I ER						
23 00 II TI 19 13 III Te 23 44 II SI 19 42 III SI	22 30 III Se 28 18 14 III ED 11 19 04 I ER 20 46 III ER						
12 01 37 11 Te 22 22 111 Se 02 24 11 Se 29 18 17 11 SI	14 20 39 II TI 30 20 06 II OD						
13 20 25 II ER 19 21 II Te 15 03 40 I TI 20 56 II Se							
04 07 I SI 31 01 38 I TI	From Move to September Junitaria antal						
16 00 59 I OD 02 26 I SI 03 41 I ER 22 57 I OD	From May to September Jupiter's satel- lites I, III, IV, are eclipsed on the west						
NOVEMBER	side of the planet, and in January and November and December on the east side. The disappearance of satellites I and II is						
d h m Sat. Phen. d h m Sat. Phen. 1 02 00 I ER 19 04 II TI	The disappearance of satellites I and II is visible May to September, and the re- appearance in January and February and October to December. Both disappearance						
20 04 I TI 20 55 II SI	October to December. Both disappearance and reappearance of satellite III are visible in						
20 55 I SI 21 41 II Te 22 17 I Te 23 33 II Se 23 08 I Se 7 17 44 II ER	January, May to August; and November and						
2 20 29 I ER 8 00 45 I OD	December; satellite IV in January and May to August. Satellite IV is eclipsed during						
4 20 03 III TI 22 51 I SI	1939.						
22 41 III Te 9 00 04 I Te 23 44 III SI 01 03 I Se							
5 02 23 III Se 19 12 I OD							

	TI	ΓΑΝ		CONJUNCTION WITH PLANET IAPETUS								
Easte	rn	Weste	rn	No	· · · · · · · · · · · · · · · · · · ·							
d	h	d	h	Inferi		Sur	perior					
Jan. 3	21.7	Jan. 11	23.9	d	h	đ	h					
" 19	21.2	27	23.5			Jan. 23						
Feb. 4	21.2			Aug. 14	08.8	July 4						
June 29	02.1	July 7	02.6	Oct. 31	16.8	Sept. 22						
July 15	01.9	23	02.0			Dec. 9	06.6					
· · · 31	01.2	Aug. 8	01.1									
Aug. 16	00.1	·· 23	23.7			TITAN						
. 31	22.5	Sept. 8	21.9	<u></u>								
Sept. 16	20.4	" 24	19.8	Inferior Superi								
Oct. 2	18.0	Oct. 10	17.4				Superior					
·· 18	15.4	·' 26	14.9	d d	h	d T 15						
Nov. 3	12.7	Nov. 11	12.4	Jan. 8	02.1	Jan. 15	19.3					
" 19	10.2	··· 27	10.1	" 24	01.7	· · · 31						
Dec. 5	08.0	Dec. 13	08.2	T 1 9	07.0	June 24						
" 21	06.2	·· 29	06.7	July 3	05.9	July 10						
				19	05.5	20						
				Aug. 4	04.7	Aug. 11						
	IAP	ETUS		20	03.4							
	1/11	6103		Sept. 5	01.7	Sept. 12						
-				20	23.6	. 28						
Easte		Weste		Oct. 6	21.1	Oct. 14						
d	h	_ d	h	22	18.5	30						
Feb. 12	00.3	Jan. 4	04.2	Nov. 7	16.0	Nov. 15						
July 24	11.2	Sept. 3	09.3	23	13.6	Dec. 1						
Oct. 11	06.7	Nov. 20	13.4	Dec. 9	11.5	17						
Dec. 28	15.2			·· 25	09.9	Jan. 2	2 02.4					

GREATEST EASTERN AND WESTERN ELONGATIONS OF SATURN'S SATELLITES TITAN AND IAPETUS

LUNAR OCCULATIONS

Prepared by J. F. HEARD

When the moon passes between the observer and a star that star is said to be occulted by the moon and the phenomenon is known as a lunar occultation. The passage of the star behind the east limb of the moon is called the immersion and its appearance from behind the west limb the emersion. As in the case of eclipses, the times of immersion and emersion and the duration of the occultation are different for different places on the earth's surface. The tables given below, adapted from the 1939 Nautical Almanac, give the times of immersion or emersion or both for occultations of stars brighter than magnitude 5.0 visible at Toronto and at Montreal and also at Vancouver and Calgary, at night. Occultations of stars fainter than magnitude 4.5 are excluded for 24 hours before and after Full Moon. Emersions at the bright limb of the moon are given only in the case of stars brighter than magnitude 3.5, and immersions at the bright limb only in the case of stars brighter than magnitude 4.5; so that most of the phenomena listed take place at the dark limb. The terms a and b are for determining corrections to the times of the phenomena for stations within 300 miles of Toronto or Montreal in the first table, and within 300 miles of Vancouver or Calgary, in the second table. Thus if λ_0 , ϕ_0 , be the longitude and latitude of the standard station and λ , ϕ , the longitude and latitude of the neighbouring station then for the neighbouring station we haveStandard Time of phenomenon = Standard Time of phenomenon at the standard station

$$+a(\lambda - \lambda_0) + b(\phi - \phi_0)$$

where $\lambda - \lambda_0$ and $\phi - \phi_0$ are expressed in degrees. The quantity *P* in the table is the position angle of the point of contact on the moon's disc reckoned from the north point towards the east. The moon occults Uranus Jan. 29th.

				I	Age	1	Toron	to	Montreal				
Date	Star		Mag.	or E*	of Moon	E.S.T.	а	b	Р	F.S.T.	a	b	Р
" 11 May 1-2 " 22 " 5 June 28 July 22 " 25 Aug. 27 Sept. 6 " 9 Nov. 15	$ \begin{split} & 5 & \text{Psc} \\ & \beta^1 & \text{Scr} \\ & \beta^1 & \text{Scr} \\ & a & \text{Vir} \\ & \rho & \text{Scr} \\ & \phi & \text{Scr} \\ & \phi & \text{Scr} \\ & \phi & \text{Scr} \\ & \phi & \text{Scr} \\ & \phi & \text{Scr} \\ & \beta & \text{Cap} \\ & \beta & \beta & \text{Cap} \\ & \beta & \beta & \beta \\ & \beta & \beta & \beta \\ & \beta & \beta &$	m. m.	$\begin{array}{c} 4.66\\ 2.99\\ 1.22\\ 3.36\\ 4.3\\ 2.22\\ 3.22\\ 4.3\\ 4.3\\ 4.3\\ 4.3\\ 4.3\\ 4.3\\ 4.3\\ 4.3$	ЕПЕЛЕЛЕНИЕНЕНЕНЕ		$\begin{array}{c} & h & m \\ Low \\ 18 & 28.1 \\ 2 & 31.6 \\ 3 & 07.2 \\ 23 & 47.5 \\ 0 & 02.3 \\ 3 & 50 \\ 0 & 02.3 \\ 3 & 50 \\ 0 & 02.3 \\ 3 & 50 \\ 0 & 02.3 \\ 3 & 50 \\ 0 & 02.3 \\ 2 & 03.4 \\ 0 & 14.7 \\ 1 & 15.7 \\ 2 & 03.1 \\ 2 & 03.1 \\ 2 & 03.2 \\ 2 & 05.1 \\ 2 & 08.3 \\ 21 & 09.2 \\ 2 & 08.3 \\ 21 & 09.2 \\ 2 & 20.8 \\ 3 & 21 \\ 2 & 20 \\ 8.3 \\ 2 & 20 \\ 8.3 \\ 1 & 25.2 \\ 2 & 20 \\ 8.3 \\ 1 & 25.2 \\ 2 & 20 \\ 8.3 \\ 1 & 25.2 \\ 2 & 20 \\ 8.3 \\ 1 & 25.2 \\ 2 & 20 \\ 8.3 \\ 1 & 25.2 \\ 2 & 20 \\ 8.3 \\ 1 & 25.2 \\ 2 & 20 \\ 8.3 \\ 1 & 25.2 \\ 2 & 20 \\ 8.3 \\ 1 & 25.2 \\ 2 & 20 \\ 8.3 \\ 1 & 25.2 \\ 2 & 20 \\ 8.3 \\ 1 & 25.2 \\ 2 & 20 \\ 8.3 \\ 1 & 25.2 \\ 2 & 20 \\ 1 $	$\begin{array}{c} -1.1\\ -2.2\\ -1.4\\ -\\ -\\ -2.0\\ -0.8\\ -1.7\\ -1.0\\ 0.0\\ -2.4\\ -1.7\\ -0.1\\ -0.1\end{array}$	$\begin{array}{c} -2.2\\ +1.8\\ -1.2\\ -\\ -\\ -0.4\\ +1.1\\ -2.0\\ +0.9\\ +2.3\\ +1.3\\ -1.5\\ -0.2\\ +1.2\\ +2.4\end{array}$	$\begin{array}{c} 334\\ 57\\ 325\\\\\\ 97\\ 29\\ 288\\ 102\\ 234\\ 264\\ 119\\ 252\\ 271\\ 59\\ \end{array}$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{c}$	$\begin{array}{c} -0.8\\ +1.7\\ +0.3\\ -2.6\\ +2.0\\ -1.8\\ -0.4\\ -0.6\\ +0.4\\ -0.6\\ +0.7\\ +2.2\\ +1.4\\ -2.0\\ +0.8\\ +2.2\\ +1.4\\ -0.8\end{array}$	$\begin{array}{r}$

LUNAR OCCULTATIONS VISIBLE AT TORONTO AND MONTREAL 1939

*Immersion or Emersion.

LUNAR OCCULTATIONS VISIBLE AT VANCOUVER AND CALGARY 1939

Date	Star	Mag.	I or	Age	N N	ancou	ver		Calgary			
Date	Star	wiag.	E*	Moon	P.S.T.	a	b	Р	M.S.T.	a	b	Р
Feb. 26 ϵ Mar. 8 a " 28-29 λ Apr. 11 ρ " 11 ρ " 1 a June 12 ϵ Sept. 5 δ " 5-6 68	Tau Vir Vir Gem Sgr Sgr Vir Vir Vir Psc Tau Tau	$\begin{array}{c} 3.6\\ 1.2\\ 1.2\\ 3.6\\ 4.0\\ 4.8\\ 1.2\\ 1.2\\ 1.2\\ 4.4\\ 3.9\\ 4.2 \end{array}$		$\begin{array}{c} d \\ 7.8 \\ 17.0 \\ 17.0 \\ 8.2 \\ 21.4 \\ 21.4 \\ 11.7 \\ 12.5 \\ 12.5 \\ 24.3 \\ 22.1 \\ 22.2 \end{array}$	h m 19 40.1 0 24.7 1 25.2 23 23.6 Low 3 30.3 1 03.3 19 40.6 20 48.6 2 34.0 Low	$ \begin{array}{r} -0.9 \\ -0.4 \\ -1.2 \\ -0.9 \\ -0.8 \\ -1.0 \end{array} $	$-0.8 \\ -1.7 \\ +0.9 \\ -2.0 \\ +0.8 \\ +0.2$	$327 \\ 109 \\ \\ 292 \\ 145 \\ 109 \\ 298 \\ 55 \\ 109 \\ 298 \\ 55 \\ 100 \\ $	$\begin{array}{c} 0 & 24.2 \\ 3 & 37.1 \\ 4 & 43.1 \\ 2 & 07.9 \\ 20 & 50.5 \\ 21 & 57.8 \end{array}$	-0.6-0.3-1.3-1.5-0.7-1.1-1.0+0.4	$\begin{array}{c} -1.6\\ -1.6\\ +2.1\\ +0.6\\ -2.0\\ +0.9\\ -0.3\\ \hline \\ +2.2 \end{array}$	$ \begin{array}{r} 343 \\ 99 \\ 48 \\ 295 \\ 138 \\ 96 \\ 311 \\ 219 \\ \end{array} $
$\begin{array}{cccc} & & 5-6 & 68 \\ & & 21 & \rho \\ Nov. & 16 & \beta \\ & & 16 & \beta \\ & & 16 & \beta \\ & & 27-28 & 119 \\ Dec. & 24 & \delta \end{array}$	Tau Sgr Cap Cap Tau Tau	$\begin{array}{r} 4.2 \\ 4.0 \\ 3.2 \\ 3.2 \\ 4.7 \\ 3.9 \end{array}$	E I I E E I	$\frac{8.8}{5.8}$	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	-1.7 -1.1 -1.0 -1.3	$-2.3 \\ -0.2 \\ -1.2 \\ +0.9$	$120 \\ 56 \\ 263 \\ 272$	Low 19 54.8 Low 0 24.5	-0.9	-0.6	63

^{*}Immersion or Emersion.

METEORS OR SHOOTING STARS

By Peter M. Millman

Meteors are small fragmentary particles of iron or stone, the debris of space, which, on entering the earth's atmosphere at high velocity, ignite and are in general completely vaporized. On a clear moonless night a single observer should see on the average about 7 meteors per hour during the first six months of the year and approximately twice this number during the second half of the year. The above figures are averages over the whole night, however, and it should be noted that meteors are considerably more numerous during the second half of the night at which time the observer is on the preceding hemisphere of the earth in its journey around the sun.

In addition to the so-called sporadic meteors there are well-marked groups of meteors which travel in elliptical orbits about the sun and appear at certain seasons of the year. The meteors of any one group, or shower, move along parallel paths and hence, owing to the laws of perspective, seem to radiate from a point in the sky known as the radiant. The shower is usually named after the constellation in which the radiant is located. The following table lists the chief meteoric showers of the year. The material was collected from different sources, including the publications of Denning and Olivier.

	Approx.	Radiant	Maximum	Hourly No. (all	Duration	Abbre-	
Shower	a	δ	Date	meteors)	(in days)	viation	
Quadrantids Lyrids Eta Aquarids Delta Aquarids Perseids Orionids Leonids Geminids	$232^{\circ} \\ 280 \\ 336 \\ 340 \\ 47 \\ 96 \\ 152 \\ 110$	$+52^{\circ}$ +37 -1 -17 +57 +15 +22 +33	Jan. 3 Apr. 21 May 4 July 28 Aug. 12 Oct. 22 Nov.16 Dec. 12	$20\\10\\10\\20\\50\\20\\20\\30$	$4 \\ 4 \\ 8 \\ 3 \\ 25 \\ 14 \\ 14 \\ 14$	Q Y E D P O L	

The Chief Annual Meteor Shower's for the Northern Hemisphere.

The date of maximum given above applies to either morning or evening and is approximate only, as local irregularities in the showers in addition to the effect of leap year may shift it by a day or more. With the exception of the Geminids, all the showers listed are most active well after midnight. It should be noted that large numbers of meteors appeared on June 28, 1916, and on Oct. 9, 1933, and there is the possibility of a return of these showers.

A meteor observer should make as complete a record as he can with efficiency. The most important information to note includes the number of meteors per hour, their magnitudes and positions in the sky, evidences of enduring trains and, where several stations are co-operating, the exact time of the appearance of each meteor. Magnitudes of meteors are generally determined by comparison with stars and the positions of meteor trails may most conveniently be recorded by plotting them as straight lines on gnomonic star maps. The observer should also make sure that the record sheet contains his name, the exact place of observation, the night when the observations were made given as a double date (e.g. the evening of May 4 or the morning of May 5 would be recorded as May 4-5), and finally, a note on the weather conditions.

The first curve shown in the figure below gives the expected hourly rate of meteors for a single observer at different times of the year. It has been drawn from data published by Denning, Olivier, and Hoffmeister. This curve varies somewhat from year to year. The corresponding curve for the southern hemisphere, which is not plotted, lacks the high maximum at P, has its highest maxima at E and D, and best general rates from April through July.

The second curve gives the number of meteor photographs found on all Harvard patrol plates up to Oct. 15, 1936, for each five-day interval throughout the year, taken from a catalogue of meteor photographs published by Miss Hoffleit. Since these plates were exposed on a uniform system the curve gives some indication of the favourable periods for meteor photography. The high photographic efficiency of the Geminid shower is a marked feature.



Of recent years the study of meteors has become increasingly important both because of its cosmic significance and because of its close association with studies of the upper atmosphere. The amateur who does not possess a telescope can render more real assistance in this field than in any other. In particular, all observations of very bright meteors or fireballs should be reported immediately in full. Maps and instructions for meteor observations may be secured from the writer at the Dunlap Observatory, Richmond Hill, Ont., the Canadian headquarters for the collection of meteor data.

For more complete instructions concerning the visual observation of meteors see the JOURNAL of the Royal Astronomical Society of Canada, vol. 31, p. 255, 1937; and for meteor photography volume 31, p. 295, 1937.

PRINCIPAL ELEMENTS OF THE SOLAR SYSTEM

Planet	Mean L from (a	Sun	Period	Eccen- tri-	In- clina-	Long. of	Long. of Peri-	Long. of	
	⊕ =1	millions of miles	(P)	city (e)	tion (i)	Node (($\begin{array}{c} \text{helion} \\ (\pi) \end{array}$	Planet	
					0	o	0	o	
Mercury	.387	36.0	88.0days	.206	7.0	47.6	76.5	96.3	
Venus,	.723	67.2	224.7	.007	3.4	76.1	130.7	259.3	
Earth	1.000	92.9	365.3	.017			101.9	99.5	
Mars	1.524	141.5	687.0	.093	1.9	49.1	334.9	7.3	
Jupiter	5.203	483.3	11.86yrs.	.048	1.3	99.8	13.3	311.8	
Saturn	9.54	886.	29.46	.056	2.5	113.1	91.8	11.5	
Uranus	19.19	1783.	84.0	.047	0.8	73.7	169.7	46.7	
Neptune	30.07	2793.	164.8	.009	1.8	131.1	44.1	168.6	
Pluto	39.46	3666.	247.7	.249	17.1	109.5	223.4	148.0	

ORBITAL ELEMENTS (Jan. 1, 0^h, 1938)

PHYSICAL ELEMENTS

Object	Symbol	Mean Dia- meter miles	Mass $\oplus = 1$	Density water =1	Axial Rotation	Mean Sur- face Grav- ity $\oplus = 1$	Albedo Bond's	tud Opj tio Elo	e at posi- n or nga- on
Sun	\odot	864,000	332,000	1.4	$24^{d}7$ (equa-	27.9		-	26.7
Moon		2,160	.0123	3.3	torial) 27 ^d 7.7 ^h	. 16	.07	-	12.6
Mercury	ĝ	3,010	.056	3.8	88 ^d	.27	.07		$0\pm$
Venus		7,580	.82	4.9	30 ^d ?	.85	. 59	-	$4\pm$
Earth	\oplus	7,918	1.00	5.5	$23^{\rm h}$ $56^{\rm m}$	1.00	.29		
Mars	ď	4,220	.108	4.0	$24^{\rm h} 37^{ m m}$. 38	.15	-	$2\pm$
Jupiter	24	87,000	318.	1.3	$9^{h} 50^{m} \pm$	2.6	. 56?	-	$2\pm$
Saturn	þ	72,000	95.	.7	$10^{b}15^{m} \pm$	1.2	.63?		$0\pm$
Uranus	ô	31,000	14.6	1.3	$10^{\rm h}.8\pm$.9	.63?	+	5.7
Neptune		33,000	17.2	1.3	16 ^h ?	1.0	.73?	+	7.6
Pluto		4,000?	<.1		-			+	14

SATELLITES OF THE SOLAR SYSTEM

Name	Stellar Mag.		Dist. from Planet Miles	d I	Perio h	d m	Diamete Miles	er Discoverer		
			[Milles							
SATELLITE										
Moon	-12.6	530	238,857	27	07	43	2160			
SATELLITES OF MARS										
Phobos	12	8	5,800	0	07	39	10?	Hall, 1877		
Deimos	13	21	14,600	1	06	18	5?	Hall, 1877		
SATELLITES	OF IUI	DITED								
V	13	48	112,600	0	11	57	100?	Barnard, 1892		
v Io	$13 \\ 5$	112^{40}	261,800	1	18	$\frac{37}{28}$	2300	Galileo, 1610		
Europa		178	416,600	$\hat{3}$	13^{10}	14^{10}	$2000 \\ 2000$	Galileo, 1610		
Ganymede	5	284	664,200	$\tilde{7}$	$\tilde{03}$	$\overline{43}$	3200	Galileo, 1610		
Callisto	$\tilde{6}$	499	1,169,000	16	16	32	3200	Galileo, 1610		
VI	14	3037	7,114,000	250	16		100?	Perrine, 1904		
VII	16	3113	7,292,000	260	01		40?	Perrine, 1905		
X	18	3116	7,300,000	260		·	15?	Nicholson, 1938		
XI	18	5990	14,000,000				15?	Nicholson, 1938		
VIII	16	6240	14,600,000				40?	Melotte, 1908		
IX	17	6360	14,900,000	758		l	20?	Nicholson, 1914		
SATELLITES	5 OF SAT	URN								
Mimas	12	27	115,000	0	22	37)	400?	W. Herschel, 1789		
Enceladus	$\overline{12}$	34	148,000	1	08	53	500?	W. Herschel, 1789		
Tethys	11	43	183,000	1	21	18	800?	G. Cassini, 1684		
Dione	11	55	234,000	2	17	41	700?	G. Cassini, 1684		
Rhea	10	76	327,000	4	12	25	1100?	G. Cassini, 1672		
Titan	8	177	759,000	15	22	41	2600?	Huygens, 1655		
Hyperion	13	214	920,000	21	06	38	300?	G. Bond, 1848		
Iapetus	11	515	2,210,000	79	07	56	1000?	G. Cassini, 1671		
Phoebe	14	1870	8,034,000	550			200?	W. Pickering, 1898		
SATELLITES	S OF UR	ANUS								
Ariel	16	14	119,000	2	12	29	600?	Lassell, 1851		
Umbriel	16	19	166,000	$\overline{4}$	$\hat{0}\bar{3}$	$\overline{28}$	400?	Lassell, 1851		
Titania	14	$\tilde{32}$	272,000	8	16	$\overline{56}$	1000?	W. Herschel, 1787		
Oberon	14	42	364,000	13	11	07	900?	W. Herschel, 1787		
SATELLITE	OF NEP	TUNE								
(Triton)		16	220,000	5	21	03	3000?	Lassell, 1846		
(Inton)	10	10	220,000	<u> </u>	41	00	50001	1010		

*As seen from the sun.

DOUBLE AND MULTIPLE STARS

By FRANK S. HOGG

A number of the stars which appear as single to the unaided eye may be separated into two or more components by field glasses or a small telescope. Such objects are spoken of as *double* or *multiple stars*. With larger telescopes pairs which are still closer together may be resolved, and it is found that, up to the limits of modern telescopes, over ten per cent. of all the stars down to the ninth magnitude are numbers of double stars.

The possibility of resolving a double star of any given separation depends on the diameter of the telescope objective. Dawes' simple formula for this relation is d''=4.5/A, where d is the separation, in seconds of arc, of a double star that can be just resolved, and A is the diameter of the objective in inches. Thus a one-inch telescope should resolve a double star with a distance of 4''.5between its components, while a ten-inch telescope should resolve a pair 0''.45 apart. It should be noted that this applies only to stars of comparable brightness. If one star is markedly brighter than its companion, the glare from the brighter makes it impossible to separate stars as close as the formula indicates. This formula may be applied to the observation of double stars to test the quality of the seeing and telescope.

It is obvious that a star may appear double in one of two ways. If the components are at quite different distances from the observer, and merely appear close together in the sky the stars form an *optical* double. If, however, they are in the same region of space, and have common proper motion, or orbital motion about one another, they form a *physical* double. An examination of the probability of stars being situated sufficiently close together in the sky to appear as double shows immediately that almost all double stars must be physical rather than optical.

Double stars which show orbital motion are of great astrophysical importance, in that a careful determination of their elliptical orbits and parallaxes furnishes a measure of the gravitational attraction between the two components, and hence the mass of the system.

In the case of many unresolvable close doubles, the orbital motion may be determined by means of the spectroscope. In still other doubles, the observer is situated in the orbital plane of the binary, and the orbital motion is shown by the fluctuations in light due to the periodic eclipsing of the components. Such doubles are designated as *spectroscopic* binaries and *eclipsing* variables.

The accompanying table provides a list of double stars, selected on account of their brightness, suitability for small telescopes, or particular astrophysical interest. The data are taken chiefly from Aitken's New General Catalogue of Double Stars, and from the Yale Catalogue of Bright Stars. Successive columns give the star, its 1900 equatorial coordinates, the magnitudes and spectral classes of its components, their separation, in seconds of arc, and the approximate distance of the double star in light years. The last column gives, for binary stars of well determined orbits, the period in years, and the mean separation of the components in astronomical units. For stars sufficiently bright to show colour differences in the telescope used, the spectral classes furnish an indication of the colour. Thus O and B stars are bluish white, A and F white, G yellow, K orange and M stars reddish.

A good reference work in the historical, general, and mathematical study of double stars is Aitken's *The Binary Stars*.

REPRESENTATIVE	DOUBLE	STAR
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S	tar	a	1900	δ		Mag. and Spect.	d	D	Remarks
α γ	And Cas UMi Ari Pis	00 01 01	$43.0 \\ 22.6 \\ 48.1$	$+57 \\ +88 \\ +18$	$17 \\ 46 \\ 48$	4.4B3; 8.5 3.6F8; 7.2M0 var. F8; 8.8 4.8A0; 4.8A0 5.2A2; 4.3A2	$"36 \\ 8 \\ 19 \\ 8.3 \\ 2.4$	${\begin{array}{c} {\rm L.Y.}\\ {\rm 410}\\ {\rm 18}\\ {\rm 270}\\ {\rm 200}\\ {\rm 162} \end{array}}$	† 479y; 66AU Polaris ††
$\begin{array}{c} \dot{6} \\ \eta \\ 32 \end{array}$	And Tri Per Eri Ori	02 02 03	$ \begin{array}{r} 06.6 \\ 43.4 \\ 49.3 \end{array} $	$^{+29}_{+55}_{-03}$	50 29 15	2.3K0; 5.4A0; 6.6 5.4G4; 7.0F3 3.9K0; 8.5 5.0A; 6.3G5 0.3B8; 7.0	$10, 0.7 \\ 3.6 \\ 28 \\ 6.7 \\ 9$	$\begin{array}{c} 220 \\ 270 \\ 360 \\ 330 \\ 540 \end{array}$	5.5y; 23AU ††
β 12 a	Ori Mon Lyn CMa Gem	06 06 06	$24.0 \\ 37.4 \\ 40.7$	$-06 \\ +59 \\ -16$	58 33 35	5.4;6.8; 6.8; 7.9; O 4.7B2; 5.2; 5.6 5.3A2; 6.2; 7.4 -1.6A0; 8.5F 3.5F0; 8.0M0	$13, 17 \\7, 25 \\1.7, 8 \\11 \\6.8$	330 190	Trapezium † 50y; 20AU †
ζ γ ξ	Gem Cnc Leo UMa Leo	08 10 11	$06.5 \\ 14.5 \\ 12.9$	$^{+17}_{+20}_{+32}$	57 21 06	$\begin{array}{c} 2.0A0; 2.8A0; 9M10\\ 5.6G0; 6.0; 6.2\\ 2.6K0; 3.8G5\\ 4.4G0; 4.9G0\\ 4.1F3; 6.8F3 \end{array}$	$\begin{array}{c} 4,70\ 1,5\ 4\ 2\ 2\ 2\ \end{array}$	71 140	340y; 79AU 60y; 21AU ††60y; 20AU
α ζ π	Vir CVn UMa Boo Boo	$\begin{array}{c} 12 \\ 13 \\ 14 \end{array}$	$51.4 \\ 19.9 \\ 36.0$	$^{+38}_{+55}_{+16}$	51 27 51	3.6F0; 3.7F0 2.9A0; 5.4A0 2.4A2; 4.0A2 4.9A0; 5.1A0 2.7K0; 5.1A0	$egin{array}{c} 6 \\ 20 \\ 14 \\ 6 \\ 3 \end{array}$	$38 \\ 130 \\ 76 \\ 200 \\ 180$	178y; 42AU †† †† †
δ ξ a	Boo Ser Sco Her Her	15 15 17	$30.0 \\ 58.9 \\ 10.1$	$+10 \\ -11 \\ +14$	52 06 30	4.8G5; 6.7 4.2F0; 5.2F0 5.1F3; 4.8; 7G7 var.M5; 5.4G 3.2A0; 8.1G2	$3 \\ 4 \\ 1, 7 \\ 5 \\ 11$	$ \begin{array}{c c} 130 \\ 86 \\ 470 \end{array} $	151y; 31AU 44.7y; 19AU † † Optical
β a γ	Lyr Cyg Cap Del Cyg	19 20 20	$26.7 \\ 12.3 \\ 42.0$	$+27 \\ -12 \\ +15$	$45 \\ 50 \\ 46$	$\begin{array}{c} 5.1,\ 6.0A3;\ 5.1,\ 5.4A5\\ 3.2K0;\ 5.4B9\\ 3.8G5;\ 4.6G0\\ 4.5G5;\ 5.5F8\\ 5.6K5;\ 6.3K5\end{array}$	$\begin{array}{c} 3, 2 \\ 34 \\ 376 \\ 10 \\ 23 \end{array}$	230 220 96 11	Pairs 207'' † Optical
ς δ	Cep Agr Cep Lac Cas	$22 \\ 22 \\ 22$	$23.7 \\ 25.5 \\ 31.4$	-00 + 57 + 39	32 54 07	var.B1; 8.0A3 4.4F2; 4.6F1 var.G0; 7.5A0 5.8B3; 6.5B5 5.1B2; 7.2B3	$14 \\ 3 \\ 41 \\ 22 \\ 3$	410 120 650 650	† †

† or ††, one, or two of the components are themselves very close visual double or, more generally, spectroscopic binaries.

VARIABLE STARS

By Frank S. Hogg

Of the naked eyes stars visible to a northern observer, nearly a hundred are known to undergo variations in their light. With field glasses or a small telescope the number of variables is enormously increased. Thus there is no dearth of material with which an inquisitive amateur may satisfy himself as to the reality and nature of the fluctuations of the light of stars. Further this curiosity may be turned to real scientific value, in that the study of variable stars is one of the best organized and most fruitful fields of research for amateur observers. For years the professional astronomer has entrusted the visual observation of many of the most important variable stars entirely to amateurs, as organized into societies in England in 1890, America in 1911, and France in 1921. The American Association of Variable Star Observers has charts of the fields of 350 of these stars, and in general supervises the work of amateur observers. The Recorder is Mr. Leon Campbell, at the Harvard Observatory, Cambridge, Massachusetts. New observers are welcomed, and supplied with charts.

In our galaxy there are already known about 5,000 variables, while in globular clusters and outside systems there are some 3,000 more. Almost all those which have been sufficiently studied may be conveniently classified, according to their light variation into ten groups, by Ludendorff's classification. His classes, with their typical stars, are listed as follows:

- I. New or temporary stars: Nova Aquilae 3, 1918.
- II. Nova-like variables: T Pyxidis, RS Ophiuchi.
- III. R Coronae stars: R. Coronae Borealis. Usually at constant maximum, with occasional sharp minima.
- IV. U Geminorum stars: U Geminorum. Usually at constant minimum, with occasional sharp maxima.
- V. Mira stars: oCeti. Range of several magnitudes, fairly regular period of from 100 to 600 days.
- VI. μ Cephei stars: μ Cephei. Red stars with irregular variations of a few tenths of a magnitude.
- VII. RV Tauri stars: RV Tauri. Usually a secondary minimum occurs between successive primary minima.
- VIII. Long period Cepheids: δCephei. Regular periods of one to forty-five days. Range about 1.5 magnitudes.
 - IX. Short period Cepheids: RR Lyrae. Regular periods less than one day. Range about a magnitude.
 - X. Eclipsing stars: β Persei. Very regular periods. Variations due to covering of one star by companion.

1939 maxima of bright variable stars (E.S.T.)

- o Ceti Aug. 31
- δ Cep Jan. 1.9, 7.5, etc.

β Lyr Jan. 5.3, Jan. 18.2, etc. R Sct Jan. 9, May 31, Oct. 18

χ Cyg Aug. 14

REPRESENTATIVE BRIGHT VARIABLE STARS

	1 .	1	1	1	1		1	1
Name	Design.	Max.	Min.	Sp.	Period	Туре	Date	Discoverer
$\begin{array}{ccc} \eta & \text{Aql} \\ \text{N} & \text{Aql} \\ \epsilon & \text{Aur} \\ \delta & \text{Cep} \end{array}$	$ \begin{array}{r} 194700 \\ 184300 \\ 045443 \\ 222557 \end{array} $	$ \begin{array}{r} 3.7 \\ -0.2 \\ 3.3 \\ 3.6 \end{array} $	$\begin{array}{r} 4.4 \\ 10.9 \\ 4.1 \\ 4.3 \end{array}$	G4 Q F5p G0	7.17652 Irr. 9833. 5.36640	VIII I X VIII	1918 1821 1784	Pigott Bower Fritsch Goodricke
δ Cep U Cep	005381	6.8	9.2	A0	2.49293	х	1880	W. Ceraski
$\begin{array}{ccc} o & \operatorname{Cet}^1 \\ \mathrm{RR} & \operatorname{Cet} \\ \mathrm{R} & \mathrm{CrB} \\ \chi & \mathrm{Cyg} \\ \mathrm{P} & \mathrm{Cyg} \end{array}$	0214 <i>03</i> 012700 154428 194632 201437a	4.2	$10.1 \\ 9.0 \\ 13.8 \\ 14.0 \\ 6.0$	M5e F0 cG0e M7e B1qk	331.8 0.55304 Irr. 412.9 Irr.	V IX III V II	1906 1795 1686	Fabricius Oppolzer Pigott Kirch Blaeu
$\begin{array}{ccc} \mathrm{SS} & \mathrm{Cyg} \\ \mathrm{XX} & \mathrm{Cyg} \\ \varsigma & \mathrm{Gem} \\ \eta & \mathrm{Gem} \\ \mathrm{R} & \mathrm{Gem} \end{array}$	213843 200158 065820 060822 070122a	$ \begin{array}{c c} 11.4\\ 3.7\\ 3.3\end{array} $	$12.0 \\ 12.1 \\ 4.1 \\ 4.2 \\ 14.3$	Pec. A cG1 M2 Se	Irr. 0.13486 10.15353 235.58 370.1	IV IX VII V V	1904 1847 1865	Wells L. Ceraski Schmidt Schmidt Hind
$\begin{array}{lll} U & Gem \\ a & Her \\ R & Hya \\ R & Leo \\ \beta & Lyr \end{array}$	074922 171014 132422 094211 184633	8.8 3.1 3.5 5.0 3.4	$13.8 \\ 3.9 \\ 10.1 \\ 10.5 \\ 4.3$	Pec. M5 M7e M7e B5e	Irr. Irr. 414.7 310.3 12.92504	IV VI V X	1795 1670 1782	Hind W. Herschel Montanari Koch Goodricke
$\begin{array}{ccc} \text{RR Lyr} \\ a & \text{Ori}^2 \\ \text{U Ori} \\ \beta & \text{Per}^3 \\ \rho & \text{Per} \end{array}$	192242 054907 054920 030140 025838	$\begin{array}{c c} 7.2 \\ 0.2 \\ 5.4 \\ 2.3 \\ 3.3 \end{array}$	$\begin{array}{r} 8.0 \\ 1.2 \\ 12.2 \\ 3.5 \\ 4.1 \end{array}$	A5 M2 M7e B8 M4	0.56685 2070.Irr. 376.9 2.86731 Irr.	IX VI V X VI	$\begin{array}{c c} 1840 \\ 1885 \\ 1669 \end{array}$	Fleming J. Herschel Gore Montanari Schmidt
R Sge R Sct λ Tau RV Tau SU Tau	$\begin{array}{c} 200916 \\ 1842o_5 \\ 035512 \\ 044126 \\ 054319 \end{array}$	$\begin{array}{c} 8.6 \\ 4.5 \\ 3.8 \\ 9.4 \\ 9.5 \end{array}$	$10.4 \\ 9.0 \\ 4.1 \\ 12.5 \\ 15.4$	cG7 K5e B3 K0 G0e	70.84 141.5 3.95294 78.60 Irr.	VII VII X VII III	1795 1848 1905	Baxendell Pigott Baxendell L. Ceraski Cannon
a UMi ⁴ N Her N Lac	$\begin{array}{c} 012288 \\ 180445 \\ 221255 \end{array}$	$2.3 \\ 1.5 \\ 2.2$	$ \begin{array}{c} 2.4 \\ 14.0 \\ \end{array} $	cF7 Q Q	3.96858 Irr. Irr.	VIII I I	1934	Hertzsprung Prentice Peltier

¹0 Cet (Mira); ²a Ori (Betelgeuse); ³B Per (Algol); ⁴a UMi (Polaris).

Most of the data in this Table are from Prager's 1936 Katalog und Ephemeriden Veranderlicher Sterne. The stars are arranged alphabetically in order of constellations. The second column, the Harvard designation, gives the 1900 position of the star. The first four figures of the designation give the hour and minute of right ascension, the last two the declination in degrees, italicised for stars south of the equator. Thus the position of the fourth star of the list, δ Cephei, is R.A. 22h 25m, Dec. +57, (222557). The remaining columns give the maximum and minimum magnitudes, spectral class, the period in days and decimals of a day, the classification on Ludendorff's system, and the discoverer and date. In the case of eclipsing stars, the spectrum is that of the brighter component.

THE DISTANCES OF THE STARS

The measurement of the distances of the stars is one of the most important problems in astronomy. Without such information it is impossible to form any idea as to the magnitude of our universe or the distribution of the various bodies in it.

The parallax of a star is the apparent change of position in the sky which the star would exhibit as one would pass from the sun to the earth at a time when the line joining earth to sun is at right angles to the line drawn to the star; or, more accurately, it is the angle subtended by the semi-major axis of the earth's orbit when viewed perpendicularly from the star. Knowing the parallax, the distance can be deduced at once.

For many years attempts were made to measure stellar parallaxes, but without success. The angle to be measured is so exceedingly small that it was lost in the unavoidable instrumental and other errors of observation. The first satisfactory results were obtained by Bessel, who in 1838, by means of a heliometer, succeeded in determining the parallax of 61 Cygni, a 6th magnitude star with a proper motion of 5'' a year. On account of this large motion the star was thought to be comparatively near to us, and such proved to be the case. At about the same time Henderson, at the Cape of Good Hope, from meridian-circle observations, deduced the parallax of Alpha Centauri to be $0^{\prime\prime}.75$. For a long time this was considered to be the nearest of all the stars in the sky, but in 1913 Innes, director of the Union Observatory, Johannesburg, South Africa, discovered a small 11th mag. star, 2° 13' from Alpha Centauri, with a large proper motion and to which, from his measurements, he assigned a parallax of 0".78. Its brightness is only 1/20,000 that of Alpha Centauri. In 1916 Barnard discovered an 11th mag. star in Ophiuchus with a proper motion of 10" per year, the greatest on record, and its parallax is about 0''.53. It is believed to be next to Alpha Centauri in distance from us.

The distances of the stars are so enormous that a very large unit has to be chosen to express them. The one generally used is the light-year, that is, the distance travelled by light in a year, or $186,000x60x60x24x365\frac{1}{2}$ miles. A star whose parallax is 1" is distant 3.26 light years; if the parallax is 0".1, the distance is 32.6 l.-y.; if the parallax is 0".27 the distance is $3.26 \div .27 = 12$ l.-y. In other words, the distance is inversely proportional to the parallax. In recent years the word *parsec* has been introduced to express the distances of the stars. A star whose distance is 1 parsec is such that its *par*-allax is 1 *sec*-ond. Thus 1 parsec is equivalent to 3.26 l.-y., 10 parsecs = 32.6 l.-y., etc.

In later times much attention has been given to the determination of parallaxes, chiefly by means of photography, and now several hundred are known with tolerable accuracy.

THE SUN'S NEIGHBOURS

By J. A. PEARCE

Through the kindness of Dr. Adriaan van Maanen, who has supplied the fundamental data, this table has been revised to contain all stars known to be nearer than five parsecs or 16.3 light-years. One star of the former table, has been discarded, and five new members have been added, making a total of forty stars in a space of 524 cubic parsecs. With the exceptions of Sirius, Procyon and Altair, all the stars are dwarfs; the list including the three white dwarfs, Sirius B, 40 Eridani B, and van Maanen's star. Forty-five per cent. of the stars are members of binary systems.

<u> </u>	(1000))	C. I	1		I T		M	
Star	$a(1900)\delta$	_Sp_	$-\frac{\mu}{\mu}$	$-\frac{\pi}{\prime\prime}$	L.y.		_ <u>M</u>	L
-	hm °′		"	"				
Sun		G0				-26.7		
Groom 34A	0 13 + 43 27	M2		0.274	11.9	8.1	10.3	.0063
Groom 34B		M5	2.85		12.1	10.7	12.9	.0006
van Maanen	0 44 + 4 55	F3	3.01	.242	13.5	12.3	14.2	.0002
τ Ceti		G7	1.92	.292	11.2	3.6		.36
ε Eri	$3\ 28 - 9\ 48$	K1	0.96	.304	10.7	3.8		.28
40 Eri A	4 11 - 7 49	K0	4.08	.213	15.3	4.5		.30
40 Eri B		A0	4.03		15.3		11.3	.0025
40 Eri C		M6	4.03	.213	15.3		12.4	.0009
Gould 5h 243		M0	8.70	.264	12.3	9.2	11.3	.0025
aCMa A		A2	1.32	.373	8.7	-1.6	1.3	
aCMa B		F0	1 32	.373	8.7		11.3	.0025
aCMi A		F4	1.24	.303	10.8	0.5	2.9	5.8
aCMi B			1.24	.303	10.8	12.5	14.9	
Groom 1618		M0	1.45	.230	14.2		8.6	
WB 10h 234		M4e	0.49	.217	15.0		10.7	
Wolf 359		M6e	4.84	.413	7.9	13.5	16.6	
	$ 10 \ 58 + 36 \ 38 $	M2	4.78	.381	8.6		10.5	.0052
Innes	$ 11 \ 12 -57 \ 02$		2.69	.339	9.6	(12.5)		.0004
aCen A		G5	3.68	.758	4.3	0.3		1.10
aCen B		K1	3.68	.758	4.3	1.7		.30
Prox. Cen		Μ	3.85	.758	4.3		15.4	
	$16\ 25 - 12\ 24$	M5	1.24	.270	12.1		11.7	
DM - 46.11540			1.06	.239	13.6	9.4	11.3	.0025
CD - 44.11909.			1.14	.215	15.2	(12.9)		.0008
AO 17415		M4	1.33	.214	15.2		10.7	.0044
Barnard	$ 17\ 53 + 4\ 25$	M5	10.30	.541	6.0		13.4	.0004
Bu 8798A		M4	2.31	290	11.2		11.5	.0021
Bu 8798B		M5	2.31	290	11.2		12.0	
aAqu		A2	0.66	. 207	15.7	0.9		
61 Cyg A	$21 \ 02 + 38 \ 15$	K8	5.27	.301	10.8	5.6	8.0	.052
61 Cyg B		M0	5.15	.301	10.8	6.3		.028
Lac 8760	$21 \ 11 - 39 \ 15$	M1	3.53	.255	12.8	6.6		.030
e Indi	21 56 - 57 12	K8	4.70	.288	11.3	4.7	7.0	.13
Kruger 60A		M3	0.87	.247	13.2		11.2	.0028
Kruger 60B		M4	0.92	.247	13.2	10.8		.0006
BD + 43.4305	$ 22 \ 42 + 43 \ 49 $	M5e	0.86	.217	15.0		11.2	.0028
Lac 9352	22 59 - 36 26	M2	6.90	.274	11.9	7.4	9.6	
Ross 248	$23 \ 36 + 43$	M6	1.82	.319	10.2	(13.8)	14.3	.0002
DM-37.15492.	23 59 - 37 51	M3	6.11	.217	15.0	8.3	10.0	.0083
	itudos in breat			a a ma a h		thora		

Note.—Magnitudes in brackets are photographic, all others are visual. A colour index of +2.0 has been taken to compute the visual absolute magnitudes of these stars. Symbols: Sp, spectrum; μ , proper motion; π , parallax; L.-y., light-year; m, apparent magnitude; M, absolute magnitude; L, luminosity compared to the sun.

THE BRIGHTEST STARS

Their magnitudes, Types, Proper Motions, Distances and Radial Velocities

By W. E. HARPER

The accompanying table contains the principal facts regarding 257 stars brighter than apparent magnitude 3.51 which it is thought may be of interest to our amateur members. The various columns should be self-explanatory but some comments may be in order.

The first column gives the name of the star and if it is preceded by the sign || such means that the star is a visual double and the combined magnitude is entered in the fourth column. Besides the 44 thus indicated there are 12 others on the list with faint companions but for these it is not thought that there is any physical connection. In the case of the 13 stars variable in light this fourth column shows their maximum and minimum magnitudes. The 20 first magnitudes stars are set up in bold face type.

In the fifth column are given the types as revised at various observatories principally at our own, but omitting the s and n designations descriptive of the line character. The annual proper motion follows in the next column and these may not necessarily be correct to the third decimal place.

The parallaxes are taken from Schlesinger's Catalogue of Bright Stars, 1930. The distance is given also in light years in the eighth column as to the lay mind that seems a fitting unit. In only one case (a Cygni) was the parallax negative and it was entered as formerly as ".005. The absolute magnitudes in the ninth column are the magnitudes the stars would have if all were at a uniform distance of 32.6 light years ($\pi = 0$."1). At that distance the sun would appear as a star of magnitude 4.8.

The radial velocities in the last column have been taken from Vol. 18 of the Lick Publications. An asterisk * following the velocity means that such is variable. In these cases the velocity of the system, if known, is given; otherwise a mean velocity for the observations to date is set down.

Of the 257 stars or star systems here listed 144 are south and 113 north of the equator. This is to be expected from the fact that the northern half of the sky includes less of the Milky Way than the southern.

The number in each spectral class, apart from the one marked peculiar, is as follows: O, 3; B, 72; A, 55; F, 22; G, 43; K, 42 and M, 19. The B-stars are intrinsically luminous and appear in this list out of all proportion to their total number. The stars of Classes A and K are by far the most numerous but the revision of types throws many originally labelled K back into the G group.

From the last column we see that 92 velocities are starred, indicating that 36 per cent. of the bright stars, or at least one in every three, are binary in character. For visual binaries the proportion has usually been listed as one in nine. Our list shows one in six but it is only natural to expect that we would observe a higher proportion among the nearby stars, such as these are on the average.

Other relationships can be established from the list if our amateur members care to study it.

$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	km./sec. -13.0* +11.4 + 5.0* +22.8
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	$\frac{\text{km./sec.}}{-13.0^*} + 11.4 + 5.0^* + 22.8$
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	$\frac{\text{km./sec.}}{-13.0^*} + 11.4 + 5.0^* + 22.8$
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$ \begin{array}{c c c c c c c c c c c c c c c c c c c $	-13.0^{*} +11.4 + 5.0* +22.8
$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$	-13.0^{*} +11.4 + 5.0* +22.8
	+ 5.0* +22.8
γ Pegasi 8 +14 38 2.9 B2 .010 .010 326 -2.1	+ 5.0* +22.8
	+22.8
β Hydri 20 -77 49 2.9 G0 2.243 141 23 3.6	-
	+74.6*
	- 7.1*
	- 3.8
	+13.1
	- 6.8
$ \beta $ Phoenicis 1 2 -47 15 3.4 G4 .042 .021 155 0.0	- 1.2
	+ 0.1
	+ 6.8
	-17.4*
	+25.7*
a Eridani 34 - 57 44 0.6 B9 .093 .045 72 -1.1	+19
	- 8.1
	- 0.6*
	+ 7.0*
	-11.7
v Arietis 2 2 +22 59 2.2 K2 .242 .040 81 0.2	-14.3
β Trianguli 4 +34 31 3.1 A6 .161 .027 121 0.2	+10.4*
o Ceti $14 - 3\ 26 1.7 - 9.6 \text{ M6e}$.239 .013 251 -2.7	+59.8*
$ \theta \text{ Eridani} 54 - 40 42 3.4 A2 .071 .022 148 0.1 $	+11.9*
	-25.7
γ Persei 58 +53 7 3.1 F9 .012 .017 192 -0.8	$+ 1.0^{*}$
ρ Persei 59 +38 27 3.4-4.2 M6 .176 .018 181 -0.3	+28.2
	+ 5.7*
a Persei 17 +49 30 1.9 F4 .041 .020 163 -1.6	- 2.4
δ Persei 36 +47 28 3.1 B5 .047 .015 217 -1.0	-10.0*
$ \eta$ Tauri $41 +23$ 48 3.0 $ B5p .053 .013 251 -1.5 $	+10.3
ζ Persei 48 +31 35 2.9 B1 .023 .006 543 -3.2	+20.9
	+16.0
	- 6. *
	+61.7
λ Tauri 55 +12 12 3.3-4.2 B3 .015 .006 543 -2.8	+13.0*
a Reticuli 4 13 -62 43 3.4 G5 .069 .022 148 0.1	+35.6

	Star	R.A. 1900	Decl 1900		Mag.	Type	Ann. Proper Motion	Parallax	Distance in Light Years	Abs. Mag.	Rad. Vel.
		h m		'			<i>"</i>	"	-		km./sec.
	Tauri		+16		1.1	K8	.205	.057	57	-0.1	+54.1
	Doradus		-55		3.5	A0p	.003		· • · •	••••	+25.6
	Orionis		+ 6	47	3.3	F5		.124	26		+24.6
	Aurigae		+33	0	2.9	K4		.021			+17.6
e	Aurigae	55	+43	41	3.4-4.1	F2	.015	.006	543	-2.8	- 4.1*
	Aurigae	50	· ·	6	3.3	B3		.012			+ 7.8
	Leporis	1	-22		3.3	K5		.026			+ 1.0
	Eridani	3			2.9	A1		.052			- 7.
	Leporis	8			3.3	A0p	1 1	.030	109		+27.7
	Aurigae	9			0.2	G1		.068	48		$+30.2^{*}$
	Orionis	10			0.3	B8p	([.006	($+23.6^{*}$
	Orionis	19	_		3.4	B0		.007			$+19.5^{*}$
-	Orionis		+ 6		1.7	B 2		.017			+18.0
· · ·	Tauri		+28		1.8	B8	1	.035	-		+ 8.0
	Leporis	24			3.0	G2		.021			-13.5
	Orionis	27			2.4	B0		.0 0 9			$+19.9^{*}$
	Leporis	28			2.7	F6		.017			+24.7
	Orionis	31	- 5	59	2.9	08		.007	,	1	$+21.5^{*}$
	Orionis	31		16	1.8	B0	.004	.008	407		+25.8
	Tauri		+21	5	3.0	B3e		.014	1		$+16.4^{*}$
	Orionis	-	- 2	0	1.8	B0		.008			+18.0
	Columbae		-34	8	2.8	B8	.040	.022	148		+34.6
-	Orionis	43	-	42	2.2	B0		.013			+20.1
	Columbae		-35	-	3.2	K0	1 1	.019			+89.4
	Orionis				0.5 - 1.1			.012			$+21.0^{*}$
	Aurigae	52	+44	56	2.1	A0p		.029			-18.1^{*}
θ.	Aurigae	53	+37	12	2.7	A1	. 106	.032	102	0.2	+28.6
	Geminorum	69			3.2-4.2		1 1	.013		-1.2	+21.4*
•	Geminorum	17	+22		3 . 2	M3		.016			+54.8
	Canis Majoris	18	-17		2.0	B1		.012		-2.6	+34.4*
	Carinae	22	-52		-0.9	F0		.016	204	-4.8	+20.5
	Geminorum		+16	29	1.9	A2	.066		69	0.3	-11.3*
	Puppis	35	-43	6	3.2	B8		.025			+28.2*
	Geminorum		+25	14	3. 2	G9	.020	.010	326		+9.9
	Geminorum	40	+13	0	3.4	F5	.230		68		+25.1
	Canis Majoris	41	-16	35	-1.6	A2	1.315	.375	9	1.3	- 7.5*
a	Pictoris	47	-61	50	3.3	A5	.271				+20.9

Star		R.A. 1900	Decl. 1900		Mag.	Type	Ann. Proper Motion	Parallax	Distance in Light Years	Abs. Mag.	Rad. Vt1.
	 h	m	0	,	·	1		"	1	J	km./sec.
au Puppis		5 47	-50	20	2.8	G8	004	.031	105	0.2	$+36.4^{*}$
$ \epsilon$ Canis Majoris		55	-28		1.6	B1		.012	1		+30.4 +27.4
ζ Geminorum		58			3.7-4.3		1 1	.0012			$+ 6.7^*$
o ² Can. Majoris		59			3.1	B5p		.001			+48.6
o Call. Majoris		09	20	TI	0.1	pob	.000	.007	400	-2.1	T40.0
δ Can. Majoris	17	′4	-26	14	2.0	G4p	.005	.010	326	-29	+34.3*
L ² Puppis	'	10	-		3.4-6.2	•	1				+53.0
π Puppis		14			2.7	K5	.012				+15.8
β Can. Minoris		22	+ 8	29	3.1	B8	1	.024			+23.
σ Puppis		26		6	3.3	M0	.192	.027	121		+88.1
a ₂ Geminorum		28	+32	6	2.0	A2	1 1	.074			+ 6.0*
a1 Geminorum		28	+32	6	2.8	A0	.209	.074	44		- 1.2*
a Can. Minoris		34	+ 5	29	0.5	F5	1.242	.310	10	2.9	- 3.0*
β Geminorum		39	+28	16	1.2	G9	.623	.110	30		+ 3.3
ξ Puppis		45	-24	37	3.5	K1	.007	.004	815	-3.5	+ 3.7*
ζ Puppis	8	0	-39	43	2.3	08	.036				-24.
ρ Puppis		3		1	2.9	F6		.016			+46.6
$ \gamma $ Velorum		6		3	2.2	OW9	.002				+35.
le Carinae		20	-59	11	1.7	K0	.032				+11.5
o Urs. Majoris		22	-61	3	3.5	G2	.166				+19.8
e Hydrae		41	+ 6	47	3.5	F9	.193	.024	136		+36.8*
δ Velorum		42	-54	20	2.0	A0	.093	.030	109		+2.2
ζ Hydrae		50	+ 6	20	3.3	G7	.101	.016	204		+22.6
ι Urs. Majoris		52	-48	26	3.1	A4	. 500	.070	47	2.3	+12.6
λ Velorum	c c	4	-43	2	2.2	K4	.022	.018	181	-1 5	+18.4
β Carinae		12		_	1.8	AO	.192				
ι Carinae			-58		2.2	FO					+13.3
a Lyncis			+34		3.3	K8	.214				+37.4
κ Velorum		19			2.6	B3	.017				+21.7*
a Hydrae		23			2.2	K4	1 1	.016			- 4.4
θ Urs. Majoris		26	-52	8	3.3	F7	1.096				+15.8
N Velorum		28	-56	36	3.0	K5	.041	1			-13.9
ϵ Leonis		40	+24	14	3.1	G0	.045	.012	272	-1.4	+ 5.1
U Carinae		45	-64	36	3.1	F0	.019	••••	••••	• • • • •	+13.6
a Leonis	10	3	+12	27	1.3	B6	.244	.055	59	0.0	+ 2.6
q Carinae		14			3.4	K5	.045				+ 8.6
$ \gamma $ Leonis			+20		2.3	G8	.347				-36.8
	1		1 = 0			1.20			-00	•.•	

Star	R.A. 1900	Decl. 1900		Mag.	Type	Ann. Proper Motion	Parallax	Distance in Light Years	Abs. Mag.	Rad. Vel.
	h m	0	1			"				km./sec.
μ Urs. Majoris	10 16	+42	0	3.2	K4	.082	.033	99	0.8	-20.3
θ Carinae	39	-63 5	52	3.0	B0	.023	.008	407	-2.4	+24. *
η Carinae	41	-59]	10	.0-7.4	Pec.	.007				-25.0
μ Velorum	42	-48 \pm	54	2.8	G5	.084	.028	116	0.1	+ 6.9
v Hydrae	45	-15 4	10	3.3	K3	.214	.033	99	0.9	-10
β Urs. Majoris	56	+56 8	55	2.4	A3	.089	.043	76		-12.1^{*}
a Urs. Majoris	58	+62 1	17	2.0	G5	.137	.030	109	-0.7	- 8.6
$oldsymbol{\psi}$ Urs. Majoris	11 4	+45	2	3.2	K0	.067	.044	74	1.4	- 3.6
δ Leonis	9	+21	4	2.6	A2	.208	.072	45		-23.2
θ Leonis	9	+15 5	59	3.4	A2	.103	.025	130	0.4	+7.8
λ Centauri	31	-62 2	28	3.3	B9	.046	.022	148	0.0	+7.9
β Leonis	44	+15	8	2.2	A2	. 507	.095	34		- 2.3
γ Urs. Majoris		+54 1	15	2.5	A0	.095	.041	79	0.6	-11.1
δ Centauri	12 3	-50 1	10	2.9	B3e	.044	.018	181	-0.8	+ 9.
e Corvi	5	-22	4	3.2	K2	.063	.027	121	-0.4	+ 4.9
δ Crucis	10	-58 1	12	3.1	B3	.051				+26.4
δ Urs. Majoris	10	+57 8	35	3.4	A0	.113	.044	74		-12.
γ Corvi	11	-16 8		2.8	B8	.159	.021	155		- 4.2*
a ¹ Crucis	21	-623	33	1.6	B1	.048	.015	217	-2.5	-12.2*
a ² Crucis	21	-62 :	32	2.1	B3		.015			+ 0.3*
γ Corvi	25	-15	58	3.1	AO		.030			+ 8.7
llδ Crucis		-56 3		1.5	M4	.270				+21.3
β Corvi	29		1	2.8	G5		.020			- 7.7
a Muscae	31			2.9	B5		.012	-		+18.
γ Centauri	36			2.4	A0		.032			- 7.5
γ Virginis	36			2.9	FO		.085			-19.6
β Muscae	40			3.3	B3		.014			+42. *
β Crucis	42		9	1.5	B1		.011			+20.0
ε Urs. Majoris		+56 3	-	1.7	A2		.045		1	-11.9*
a Can. Venat.		+38 3		2.8	A1	1	.025			- 3.6*
ϵ Virginis		+11 3		3.0	G6		.034			-14.0
γ Hydrae	13 13	-22 3	39	3.3	G7	.085	.017	192	-0.5	- 5.4
i Centauri	15	F		2.9	A2	.351				
ζ ¹ Urs. Majoris		+552		2.4	A2p		.043	76		- 9.9*
a Virginis	20	1		1.2	B2		.017			$+ 1.6^{*}$
ζ Virginis	30		5	3.4	A2		.036			-13.1

						_			
Star	R.A. 1900	Decl. 1900	Mag.	Type	Ann. Proper Motion	Parallax	Distance in Light Years	Abs. Mag.	Rad. Vel
	h m	0 /	1		111	"	1	1	km./sec.
e Centauri	13 34	1 1	2.6	B2	040	.013	951	1 0	+ 5.6
				B2 B3				1	
η Urs. Majoris		+49 49	1.9		1 1	.013	1		-10.9
μ Centauri	44		3.3	B3e	.030		1		+12.6
5 Centauri		-46 48	3.1	B3		.010	1	-1.9	
η Boötis	50	+1854	2.8	G1		.100		2.8	
β Centauri	57	-5953	0.9	B 3	.039	.020	163	-2.6	+12.0*
π Hydrae	14 1	-26 12	3.5	K3	.165	.036	91	-1.3	+27.2
θ Centauri	1	$-35\ 53$	2.3	G8	.748	.067	49	-1.4	+ 1.3
a Boötis	11	+19 42	0.2	K0	2.287				- 5.1
γ Boötis	28	1.	3.0	A3	.182				-35.5
n Centauri	29		2.6	B3e	.052				- 0.2
la Centauri	33		0.1	G0	3.682				-22.2
a Circini	34	1.	3.4	F0		.070			+7.4
a Lupi	35		2.9	B2		.009		}	$+7.3^{*}$
· •	41		$2.9 \\ 2.7$	G8	1			1	
le Boötis						.018			+16.4
a ² Librae	45		2.9	F1		.073			-10. *
β Urs. Minoris	51	-	2.2	K4	.028	.035			+16.9
βLupi	52	1 .	2.8	B3	1	.012			- 0.3*
к Centauri	53		3.4	B2	.037	.009			+ 9.1*
σ Librae	58	-24 53	3.4	M4	.094	.024	136	0.3	- 4.3
ζ Lupi	15 5	-51 43	3.5	G5	.132	.017	192	-0.4	- 9.7
γT Australis	10	-68 19	3.1	A0	.064				0.
β Librae	12	-91	2.7	B8	.108	.024	136	-0.4	-37. *
δLupi	15	-40 17	3.4	B3	.032	.010	326	-1.6	+ 1.6
γ Urs. Minoris	21	1	3.1	A2		.042	1		- 3.9*
L Draconis	23		3.5	K3		.031			-11.1
lly Lupi	28		3.0	B3		.016			+ 6.
a Cor. Borealis	30		2.3	A0		.044			$+ 1.0^{*}$
a Serpentis	39	1.	2.8	K3		.011			+ 3.0
β T Australis	46		3.0	F0			1		- 0.3
	1			-		.090	1		
π Scorpii	53		3.0	B3		.012			- 3.0*
δ Scorpii	54	-22 20	2.5	B1	.042	.011	296	-2.3	-16. *
\$ Scorpii	16 0	-19 32	2.8	B3	.041	.005	652	-1.4	- 9.3*
δ Ophiuchi	9	- 3 26	3.3	K8	.159	.029	112	0.4	-19.8
e Ophiuchi	13	-427	3.3	G9	.088	.030	109	0.7	-10.3
σ Scorpii	15	-25 21	3.1	B1	.033	.007	466		- 0.4*
η Draconis	23	+61 44	2.9	G5	.062	.038	86	0.8	-14.3
····							1	1	

Star	R.A. 1900	Decl. 1900	Mag.	Type	Ann. Proper Motion	Parallax	Distance in Light Years	Abs. Mag.	Rad. Vel.
· · · · · · · · · · · · · · · · · · ·	lh n	· · · /	•	1		11	1	1	km. sec.
	h n	La se se		3.63	000	000	100	0.0	
lla Scorpii	16 23	1	1	M1	.032		163	-2.3	1
β Herculis	26	ſ.	ſ	G4	1 1	.021	1		-25.8*
au Scorpii	30	$ -28 \ 1$	2.9	B1	.042	.007	466		+ 0.6
ζ Ophiuchi	32	2 - 10 22	2.7	B0	.024	.009	362	-2.5	-19. *
ζ Herculis	38	1	3.0	GO	.601	.106	31	3.1	
aT Australis	38		1.9	K5	.034	.030	109	-0.7	
ε Scorpii	44	1 .	1	G9	.668		1	0.4	
•	45			B3	.032		1	-1.5	*
μ^1 Scorpii			1						80
ζ Arae	50			K5	.047		1		- 6.0
κ Ophiuchi	53	+932	3.4	K3	.296	.037	88	1.3	-55.6
η Ophiuchi	17 5	-15 36	2.6	A2	.094	.036	91	0.4	- 1.0
η Scorpii	5	-43 6	3.4	A7	.294	.069	47	2.6	-28.4
ζ Draconis	8	$+65\ 50$	3.2	B8	.023	.026	125	0.3	-14.1
a Herculis	10		3.1-3.9		.030				-32.5
δ Herculis	11	1	1	A2	.164				-39. *
		1.		K3	.021	.022			
π Herculis	12	1.	1				1	0.1	
θ Ophiuchi	16		1	B2	.030			-1.9	
β Arae	17	1		K1	.035	.017		-1.0	
v Scorpii	24	-37 13	2.8	B3	.040	.010	326	-2.2	+18. *
a Arae	24	-49 48	3.0	B3e	.085	.017	192	-0.9	
λ Scorpii	27	-37 2	1.7	B2	.040	.016	204	-2.3	0. *
β Draconis	28	$+52\ 23$	3.0	GO	.012	.008	407	-2.5	-20.1
θ Scorpii	30		1	FO	.010				+1.4
a Ophiuchi	30		1	A0	.264	.052	1	0.7	
κ Scorpii	36	1.		B3	.032				-10. *
•	38			K2	.052	-	1	1	-11.9
β Ophiuchi						.036			1
l ¹ Scorpii	41			F8	.004	.007			-27.6
$ \mu $ Herculis	43	+27 47	3.5	G5	.817	.112	1	3.7	-16.1
G Scorpii	43			K2	.068	.028	116		+24.7
v Ophiuchi	54	-946	3.5	G7	.118	. 02 3	142	0.3	+12.4
γ Draconis	54	+51 30	2.4	K5	.026	.028	116	-0.3	-27.8
γ Sagittarii	59		1	KO	.206	.041	79		$+22.3^{*}$
1									•
η Sagittarii	18 11	-36 48	3.2	M4	.223	.032	102	07	+ 0.5
, 0	1	1	1	K4	.042	.032	1		-20.0
δ Sagittarii		1		1					
η Serpentis	16	1	1	G9	.898	.060	54	2.3	+ 8.9
e Sagittarii	18		1	A0	.139	• • • •			-10.8
λ Sagittarii	22			K1	.197	.048		-1.4	
a Lyrae	34	+38 41	0.1	A1	.348	.1 2 3	26	0.6	-13.8
		·	<u>.</u>						·
Star	R.A. 1900	Decl. 1900	Mag.	Type	Ann. Proper Motion	Parallax	Distance in Light Years	Abs. Mag.	Rad. Vel.
---	---	--	--	--	--	--	---	--	--
φ Sagittarii β Lyrae σ Sagittarii γ Lyrae ζ Sagittarii	h m 18 39 46 49 55 56	$\begin{array}{rrr} -27 & 6 \\ +33 & 15 \\ -26 & 25 \\ +32 & 33 \end{array}$	3.33.4-4.12.13.32.7	B8 B2p B3 B9p A2	.011 .081 .010	" .018 .003 .018 .016 .036	1086 181 204	-4.2 - 1.6 - 0.7	km./sec. +21.5* -19.0* -10.7 -21.5* +22.1
τ Sagittarii ζ Aquilae π Sagittarii δ Draconis δ Aquilae $ \beta$ Cygni γ Aquilae $ \delta$ Cygni a Aquilae	21 27 42 42	+13 43 -21 11 +67 29 + 2 55	3.4 3.0 3.2 3.4 3.2 2.8 3.0 0.9	K0 A0 F2 G8 A3 K0 K3 A1 A2	.265 .103 .041 .135 .267 .010 .018 .067 .659	.037 .022 .032 .057 .020 .023 .034	88 148 102 57 163	$0.9 \\ -0.3 \\ 0.8 \\ 2.2 \\ 0.3 \\ -0.4 \\ 0.6$	$+45.4^{*}$ -25. * -9.8 +24.8 -32.3^{*} -23.9^{*} -2.0 -20. -26.1
 θ Aquilae β Capricorni a Pavonis γ Cygni a Indi a Cygni ϵ Cygni 	$\begin{array}{ccc} 20 & 6 \\ & 15 \\ & 18 \\ & 19 \\ & 31 \\ & 38 \\ & 42 \end{array}$	$\begin{array}{rrrr} -15 & 6 \\ -57 & 3 \\ +39 & 56 \\ -47 & 38 \\ +44 & 55 \end{array}$	$\begin{array}{c} 3.4 \\ 3.2 \\ 2.1 \\ 2.3 \\ 3.2 \\ 1.3 \\ 2.6 \end{array}$	A0 F8 B3 F8 G2 A2p G7	.035 .042 .090 .006 .072 .004 .485	.017 .013 .007 .036 .005	$\begin{array}{c} 192 \\ 251 \end{array}$	-0.6 -2.3 -3.4 1.0 -5.2	$\begin{array}{r} -28.6^{*} \\ -19.0^{*} \\ + 1.8^{*} \\ - 7.6 \\ - 1.1 \\ - 6.3^{*} \\ -10.5^{*} \end{array}$
 ζ Cygni a Cephei β Aquarii β Cephei ε Pegasi δ Capricorni γ Gruis 	16 26 27	$+70 7 \\ + 9 25$	3.4 2.6 3.1 3.3 2.5 3.0 3.2	G6 A2 G1 B1 K2 A3 B8	.061 .163 .020 .013 .028 .395 .108	.078 .006 .008 .020 .095	42 543 407 163 34	$2.1 \\ -3.0 \\ -2.2 \\ -1.0$	$+16.9^{*}$ - 8. + 6.7 - 7.2^{*} + 5.2 - 6.4^{*} - 2.1
 a Aquarii a Gruis a Tucanae β Gruis η Pegasi a P Australis β Pegasi a Pegasi γ Cephei 	38 52 59	$\begin{array}{cccc} - & 0 & 48 \\ -47 & 27 \\ -60 & 45 \\ -47 & 24 \\ +29 & 42 \\ -30 & 9 \\ +27 & 32 \\ +14 & 40 \\ +77 & 4 \end{array}$	2.92.23.11.32.62.6	G0 B5 K5 M6 G1 A3 M3 A0 K1	.200 .085 .132 .039 .367 .235	.007 .028 .023 .015 .013 .122 .020 .034	116 - 142 - 217 - 251 - 26	$ \begin{array}{r} -0.6 \\ -0.3 \\ -1.9 \\ -1.3 \\ 1.7 \\ -0.9 \\ 0.2 \\ \end{array} $	+7.6 +11.8 +42.2* +1.6 +4.4* +6.5 +8.6 -4.* -42.0

STAR CLUSTERS AND NEBULAE

Prepared by J. F. HEARD

The amateur who possesses a telescope will find great interest in the observation and identification of star clusters and nebulae. Such objects, of course, have been extensively catalogued and classified. The most frequently quoted catalogue is Dreyer's New General Catalogue (N.G.C.) containing 7,840 objects, extended by the Index Catalogue (I.C.) containing 5,386 more. The most interesting catalogue historically, however, and one which is still quoted for reference to the more conspicuous objects is Messier's Catalogue (M) which contains 103 objects. It was drawn up in 1781 by Charles Messier for his own convenience in identifying comets.

Messier's Catalogue as given below is adapted from a publication by Shapley and Davis (Pub. A.S.P., XXIX, 178, 1917). It includes the Messier number, the N.G.C. number, the 1900 position, the classification of the object and, under remarks, the name of the object (if any).

The classification is not that of Messier; it is the new classification based on modern knowledge of these objects. The clusters are classified as open clusters, which are loose irregular aggregates usually of a few scores of stars, or as globular clusters which are compact aggregates of upwards to hundreds of thousands of stars in spherical formation. The nebulae are classified as diffuse, planetary or spiral. The diffuse nebulae are great clouds of gas and "star-dust" rendered luminous by nearby stars and the planetaries are compact atmospheres of the same materials surrounding a single star. The spirals, on the other hand, are self-luminous and quite outside our stellar system and must be thought of as island universes or other galaxies like our own.

Messier	N.G.C.	R.A. (1900)	Dec. (1900)	Type of Object	Remarks
1	1952	$\begin{array}{cc} \mathrm{h} & \mathrm{m} \\ 5 & 28.5 \end{array}$	+21 57	Diffuse nebula	The Crab nebula in Taurus
$\frac{2}{3}$	7089	$21 \ 28.3$	- 1 16	Globular cluster	
	5272	13 37.6	+2853	Globular cluster	
4	6121	16 17.5	$-26\ 17$	Globular cluster	
$\frac{4}{5}$	5904	15 13.5	+227	Globular cluster	
6	6405	17 33.5	-32 9	Open cluster	
7	6475	17 47.3	-34 47	Open cluster	
8	6523	17 57.6	-24 23	Diffuse nebula	The Lagoon nebula —verv large
9	6333	17 13.3	-18 25	Globular cluster	
10	6254	$16 \ 51.9$	- 3 57	Globular cluster	
11	6705	18 45.7	-623	Open cluster	
12	6218	16 42.0	- 1 46	Globular cluster	
13	6205	16 38.1	+36 39	Globular cluster	The Hercules cluster —best example

MESSIER'S CATALOGUE OF CLUSTERS AND NEBULAE

MESSIER'S CATALOGUE OF CLUSTERS AND NEBULAE-continued

			Dec.		
Messier	N.G.C.	(1900)	(1900)	Type of Object	Remarks
14	6402	h m 1732.4	-311	Globular cluster	
15	7078	$21 \ 25.2$	+11 44	Globular cluster	
16	6611	18 13.2	-13 49	Open cluster	
17	6618	18 15.0	-16 13	Diffuse nebula	The Horseshoe or Omega nebula— bright
18	6613	18 14.1	-17 10	Open cluster	blight
$\begin{array}{c} 19 \\ 20 \end{array}$	$\begin{array}{r} 6273 \\ 6514 \end{array}$	16 56.4	$egin{array}{c c} -26 & 7 \\ -23 & 2 \end{array}$	Globular cluster	$T_{1} = T_{1} = 1$
		17 56.3		Diffuse nebula	The Trifid nebula— bright
21	6531	17 58.6	$-22\ 30$	Open cluster	
$\frac{22}{23}$	$\begin{array}{r} 6656 \\ 6494 \end{array}$	$\begin{array}{c} 18 \ \ 30.3 \\ 17 \ \ 51.0 \end{array}$	$ \begin{array}{rrrr} -23 & 59 \\ -19 & 0 \end{array} $	Globular cluster	
$\frac{23}{24}$	6603	$17 \ 51.0$ $18 \ 12.6$	$-19 \ 0$ $-18 \ 27$	Open cluster Open cluster	
25	I.C. 4725	$18 \ 12.0 \ 18 \ 25.8$	-19 19	Open cluster	
$\overline{26}$	6694	$18 \ 39.8$	-930	Open cluster	
$\overline{27}$	6853	19 55.3	+22 27	Planetary ne- bula	The Dumb-bell ne- bula
28	6626	$18 \ 18.4$	-24 55	Globular cluster	
29	6913	$20 \ 20.3$	+38 12	Open cluster	
30 21	7099	$21 \ 34.7$	-23 38	Globular cluster	The Andrewski
31	224	0 37.3	+40 43	Spiral nebula	The Andromeda ne- bula-largest
32	221	0 37.2	+40 19	Spiral nebula	spiral Very close to M31 much smaller
33	598	1 28.2	+30 9	Spiral nebula	much smaller
34	1039	$2 \ 35.6$	+42 21	Open cluster	
35	2168	6 2.7	+24 21	Open cluster	
36	1960	529.5	+34 4	Open cluster	
37	2099	5 45.8	+32 31	Open cluster	
38 39	$\begin{array}{c}1912\\7092\end{array}$	$5 \ 22.0 \\ 21 \ 28.6$	$^{+35}_{+48}$ $^{45}_{0}$	Open cluster	
40	1092	12 17.4	$+48 & 0 \\ +58 & 40$	Open cluster	Two faint stars mis-
10		12 17.4	-10 H	•••••	taken for a nebula by Messier
41	2287	6 42.7	-20 38	Open cluster	-
42	1976	5 30.4	- 5 27	Diffuse nebula	The Orion nebula
$\begin{array}{c} 43\\ 44 \end{array}$	$\begin{array}{c}1982\\2632\end{array}$	$\begin{array}{c c} 5 & 30.6 \\ 8 & 34.3 \end{array}$	-520	Diffuse nebula	Duranana au tha Daa
44	2032	8 94.9	+20 20	Open cluster	Praesepe or the Bee- hive cluster
45		$3 \ 41.5$	+23 48	Open cluster	The Pleiades
46	2437	$7 \ 37.2$	-14 35	Open cluster	
47	2478	$\begin{bmatrix} 7 & 50.2 \\ 0 & 0 \end{bmatrix}$	$ \begin{array}{ccc} -15 & 9 \\ - & 1 & 39 \end{array} $	Open cluster	
$\begin{array}{c} 48\\ 49 \end{array}$	4472	$\begin{array}{c c} 8 & 9.0 \\ 12 & 24.7 \end{array}$	-139 +833	Open cluster	
49 50	$\frac{4472}{2323}$	$\begin{bmatrix} 12 & 24.7 \\ 6 & 58.2 \end{bmatrix}$	+ 8 33 - 8 12	Spiral nebula Open cluster	
51	5194	$13 \ 25.7$	+47 43	Spiral nebula	The Whirlpool ne-
50	705 4				bula
$52 \\ 53$	$\begin{array}{c} 7654 \\ 5024 \end{array}$	$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	$+61 \ 3 \\ +18 \ 42$	Open cluster Globular cluster	
$53 \\ 54$	6715	18 48.7	$+18 42 \\ -30 36$	Globular cluster	
	0110	10 10.1	00 00	Giobulai ciustel	

MESSIER'S CATALOGUE OF CLUSTERS AND NEBULAE-continued

Messier	N.G.C.	R.A. (1900)	Dec. (1900)	Type of Object	Remarks					
		h m	0 /							
55	6809	$19 \ 33.7$	-31 10	Globular cluster						
56	6779	$19 \ 12.7$	+30 0	Globular cluster						
57	6720	$18 \ 49.9$	+3254	Planetary ne-	The Ring nebula in					
58	4579	12 32.7	+12 22	bula Spiral nebula	Lyra					
58 59	4621	$12 \ 32.1$ $12 \ 37.0$	$^{+12}_{+12}$ $^{22}_{12}$	Spiral nebula						
60	4649	$12 \ 31.0 \ 12 \ 38.6$	+12 12 +12 6	Spiral nebula						
61	4303	$12 \ 10.8$	+12 + 5 + 2	Spiral nebula						
62	6266	16 54.8	-29 58	Globular cluster						
63	5055	$10 01.0 \\ 13 11.3$	+42 34	Spiral nebula						
64	4826	12 51.8	+22 13	Spiral nebula						
65	3623	11 13.7	+13 38	Spiral nebula						
66	3627	11 15.0	+13 32	Spiral nebula						
67	2682	8 45.8	+12 11	Open cluster						
68	4590	$12 \ 34.2$	-26 12	Globular cluster						
69	6637	18 24.8	-32 $\overline{25}$	Globular cluster						
70	6681	18 36.7	-32 $\overline{23}$	Globular cluster						
71	6838	19 49.3	+18 31	Open cluster						
$\dot{72}$	6981	$20 \ 48.0$	-12 55	Globular cluster						
$\overline{73}$	6994	20 53.5	-13 1	Open cluster						
$\overline{74}$	628	$1 \ 31.3$	+15 16	Spiral nebula						
$\hat{75}$	6864	20 0.2	-22 12	Globular cluster						
76	650	1 36.0	+51 4	Planetary ne-						
			0.00	bula						
77	1068	$2\ 37.6$	-0.26	Spiral nebula						
78	2068	5 41.6	$+ 0 1 \\ -24 37$	Diffuse nebula Globular cluster						
79	1904	$\begin{array}{c}5 \hspace{0.1cm} 20.1 \\16 \hspace{0.1cm} 11.1 \end{array}$	-24 37 -22 44	Globular cluster						
80	6093 3031	9 47.3	+69 32	Spiral nebula						
$\begin{array}{c} 81 \\ 82 \end{array}$	3034	947.5 947.5	+70 10	Spiral nebula						
83	5236	$13 \ 31.4$	$-29\ 21$	Spiral nebula						
84 84	4374		+13 26	Spiral nebula						
85	4382	12 20.0 12 20.4	+18 45	Spiral nebula						
86	4406		+13 30	Spiral nebula						
80 87	4486	12 25.8	+1257	Spiral nebula						
88	4501	12 26.9	+1458	Spiral nebula						
89	4552	$12 \ 30.6$	+13 6	Spiral nebula						
90	4569	$12 \ 31.8$	+13 43	Spiral nebula						
91		$12 \ 36.0$	+13 50		Not confirmed—					
01			•		probably comet					
92	6341	17 14.1	+43 15	Globular cluster	-					
93	2447	7 40.5	-23 38	Open cluster						
94	4736		+41 40	Spiral nebula						
95	3351	10 38.7	+12 14	Spiral nebula						
96	3368		+12 21	Spiral nebula						
97	3587	11 9.0	+55 34	Planetary ne- bula	The Owl nebula					
98	4192	12 8.7	+15 27	Spiral nebula						
99	4152		+1458	Spiral nebula						
100	4321	12 17.9	+1623	Spiral nebula	1					
101	5457	13 59.6	+54 50	Spiral nebula						
$101 \\ 102$	5866?	15 3.8	+56 9	Spiral nebula						
103	581		$+60\ 11$	Open cluster						



Mi	dnig	ht	 	•	 • •	.Feb.	6
11	p.m		 		 	. "	21
10	"		 		 	. Mar.	7
9			 		 	. "	22
0			 		 	Apr.	6
7						. "	



Mi	dnig	h	t.									.May 8
11	p.m.											. " 24
10	" "	•										.June 7
9	"			•								. " 22
8	" "	•	•	•	•	•	•	•	•	•	•	. July 6



Μ	idnig	ht	• •				.Aug. 5
11	p.m						. " 21
							Sept. 7
9	" "						. " 23
8	"						. Oct. 10
7	**						. " 26
6	" "						. Nov. 6
5	**						. " 21



Mi	idnig	h	t.						. Nov.	6
11	p.m.				•				. "	21
10	"								. Dec.	6
9	" "		•						. "	21
8	"								. Jan.	5
$\overline{7}$	"								. "	20
6	"		•					•	.Feb.	6

BEGINNING OF MORNING AND ENDING OF EVENING TWILIGHT

			100 C		
	Latitude 35°	Latitude 40°	Latitude 45°	Latitude 50°	Latitude 52°
	Morn. Eve.				
Jan. 1 11 21 31 Feb. 10	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$
20 Mar. 2 12 22 Apr. 1	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$
11 21 May 1 11 21	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$
31 June 10 20 30 July 10	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	0 23 11 42	
20 30 Aug. 9 19 29	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$ \begin{array}{c ccccccccccccccccccccccccccccccccccc$
Sept. 8 18 28 Oct. 8 18	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$
Nov. 28 17 27 Dec. 7	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$
17 27 Jan. 1	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$

The above table gives the local mean time of the beginning of morning twilight, and of the ending of evening twilight, for various latitudes. To obtain the corresponding standard time, the method used is the same as for correcting the sunrise and sunset tables, as described on page 10. The entry — in the above table indicates that at such dates and latitudes, twilight lasts all night. This table, taken from the American Ephemeris, is computed for *astronomical* twilight, i.e., for the time at which the sun is 108° from the zenith (or 18° below the horizon).

TEMPERATURE AND PRECIPITATION AT CANADIAN AND UNITED STATES STATIONS

	Mean Temperature, Fahrenheit.													Averag Annua		
Station.	Jan.	Feb.	Ma	Ap.	May	Ju.	Jul.	Aug.	Sep.	Oc.	No.	De.	M	H	L.	
Victoria, B.C Vancouver, B.C Edmonton, Alta	30 36 6	$40 \\ 39 \\ 12$	44 43 22	49 48 40	$53 \\ 53 \\ 51$	57 60 57		$56 \\ 63 \\ 59$	56 57 50	$51 \\ 50 \\ 41$	$45 \\ 43 \\ 26$	41 38 14	49 50 37	86 86 89 -	$19 \\ 13 \\ -41$	
Calgary, Alta Regina, Sask Winnipeg, Man	$ \begin{array}{c} 11 \\ -4 \\ -3 \end{array} $	$\stackrel{14}{\stackrel{-2}{_2}}$	$25 \\ 14 \\ 16$	40 37 38	$49 \\ 50 \\ 52$	$56 \\ 59 \\ 62$	${61 \\ 64 \\ 62 }$	$59 \\ 61 \\ 64$	$50 \\ 51 \\ 54$	$42 \\ 39 \\ 41$	$26 \\ 21 \\ 22$	$\begin{array}{c} 20 \\ 8 \\ 6 \end{array}$	38 33 35	91 - 94 - 94 -	-40	
Toronto, Ont Ottawa, Ont Montreal, Que	$23 \\ 12 \\ 14$	$22 \\ 13 \\ 15$	$30 \\ 25 \\ 26$	$42 \\ 42 \\ 41$	$53 \\ 55 \\ 55$	${}^{63}_{65}_{65}$	69 69 70	67 66 67	60 59 59	48 46 47	37 33 33	27 17 20	$\begin{array}{c} 45\\ 42\\ 43\end{array}$	92 - 93 - 90 -	-24	
Halifax, N.S Churchill, Man Aklavik, N.W.T	-19	-17		$39 \\ 15 \\ 8$	$49 \\ 29 \\ 31$	$58 \\ 42 \\ 49$	$\begin{array}{c} 65 \\ 53 \\ 56 \end{array}$	${}^{64}_{52}_{50}$	$58 \\ 41 \\ 38$	$49 \\ 26 \\ 19$	39 7 - 4 -	-10	$44 \\ 18 \\ 16$	89 81 - 83 -	-46	
St. John's, Nfld New York, N.Y Washington, D.C	$23 \\ 31 \\ 33$	$22 \\ 31 \\ 35$	$28 \\ 37 \\ 42$	$35 \\ 49 \\ 53$	$43 \\ 60 \\ 64$	$51 \\ 68 \\ 72$	$59 \\ 73 \\ 76$	$ \begin{array}{r} 60 \\ 73 \\ 75 \end{array} $	$54 \\ 56 \\ 68$	$45 \\ 56 \\ 57$	$37 \\ 44 \\ 45$	$29 \\ 35 \\ 36$	$41 \\ 52 \\ 55$	83 95 98	$^{-6}_{2}_{4}$	
Chicago, Ill Denver, Colo San Francisco	$25 \\ 29 \\ 50$	$28 \\ 32 \\ 51$	$36 \\ 39 \\ 53$	$48 \\ 47 \\ 54$	59 57 56	${68 \atop 67 \atop 57}$	$74 \\ 72 \\ 57$	$73 \\ 71 \\ 58$	$\begin{array}{c} 66 \\ 63 \\ 60 \end{array}$	$55 \\ 51 \\ 59$	$\frac{41}{39}\\55$	$30 \\ 32 \\ 51$	$50 \\ 50 \\ 55 \\ 55$	95 - 97 - 91		

Prepared by Andrew Thomson.

M, H and L are the mean and the averages of the highest and of the lowest temperatures each year at the station, over the total time since the station was installed.

ar yana kanalar ini ing munya diserupa kanalar kanalar ini ini ini ini ini ini ini ini ini in	Mean Precipitation.					(Unit = one tenth of an inch)							Year.		
Station	Jan.	Feb.	Ma.	Ap.	May	Ju.	Jul.	Aug.	Sep.	Oc.	No.	De.	М	W	D
Victoria, B.C Vancouver, B.C Edmonton, Alta	$\begin{smallmatrix} 45\\88\\9 \end{smallmatrix}$	30 57 7	$23 \\ 52 \\ 7$	$\begin{array}{c}12\\32\\9\end{array}$	10 28 17	$9 \\ 23 \\ 31$	$\begin{smallmatrix}&4\\13\\33\end{smallmatrix}$	$\begin{array}{c} 6\\16\\24\end{array}$	$ \begin{array}{r} 15 \\ 38 \\ 13 \end{array} $	$28 \\ 58 \\ 7$	$\begin{array}{c} 43\\85\\7\end{array}$		575	$510 \\ 676 \\ 278$	
Calgary, Alta Regina, Sask Winnipeg, Man	$5 \\ 4 \\ 9$	${6 \atop 3 \atop 8}$	$\begin{array}{c} 7\\5\\11\end{array}$	$7 \\ 7 \\ 13$	$24 \\ 20 \\ 22$	${32 \atop {32} \atop {31}}$	$26 \\ 25 \\ 31$	$27 \\ 19 \\ 23$	$13 \\ 12 \\ 23$	$\begin{smallmatrix}&6\\&7\\15\end{smallmatrix}$	$\begin{array}{c} 7\\5\\11\end{array}$	4	141	$346 \\ 272 \\ 302$	79 101 102
Toronto, Ont Ottawa, Ont Montreal, Que	28 30 37	$25 \\ 25 \\ 32$	$25 \\ 26 \\ 35$	$25 \\ 22 \\ 25$	$29 \\ 28 \\ 30$	$27 \\ 32 \\ 35$	30 33 37	29 30 35	$30 \\ 27 \\ 35$	$24 \\ 28 \\ 33$	$28 \\ 25 \\ 35$	29		$436 \\ 444 \\ 530$	232
Halifax, N.S Churchill, Man Aklavik, N.W.T	$56 \\ 6 \\ 7$	$\begin{array}{c} 45\\10\\8\end{array}$	$\begin{array}{c} 50\\11\\6\end{array}$	$45 \\ 10 \\ 7$	$\substack{42\\10\\8}$	$37 \\ 20 \\ 7$	$39 \\ 18 \\ 16$	$45 \\ 25 \\ 14$	$36 \\ 26 \\ 10$	$53 \\ 13 \\ 8$	$54 \\ 12 \\ 10$	9	168	678 150	388 98
St. John's, Nfld New York, N.Y Washington, D.C	$54 \\ 36 \\ 35$	$51 \\ 41 \\ 35$	$45 \\ 35 \\ 37$	42 33 33	$36 \\ 32 \\ 36$	${36 \\ 34 \\ 42}$	$37 \\ 42 \\ 46$	$36 \\ 43 \\ 39$	$38 \\ 34 \\ 33$	$54 \\ 35 \\ 28$	${61 \\ 30 \\ 24}$	35	$538 \\ 430 \\ 422$		427 331 307
Chicago, Ill Denver, Colo San Francisco	$\begin{array}{c} 19\\ 4\\ 44 \end{array}$	$\begin{array}{c}23\\6\\42\end{array}$	$26 \\ 10 \\ 31$	$28 \\ 21 \\ 17$	$\begin{array}{c} 35\\22\\8\end{array}$	$\begin{array}{c} 34\\14\\2 \end{array}$	$\begin{smallmatrix} 33\\17\\0\end{smallmatrix}$	$\begin{array}{c} 32\\14\\0\end{array}$	$\substack{\substack{32\\10\\4}}$	$25 \\ 11 \\ 11 \\ 11$	$\begin{array}{c}24\\6\\24\end{array}$	7	$327 \\ 141 \\ 220$		$244 \\ 79 \\ 91$

M, W and D indicate the mean, the greatest and the least total precipitation in one year from Jan. 1 to Dec. 31 recorded at a station, records being available for varying periods from 30 to 50 years.



THE ROYAL ASTRONOMICAL SOCIETY OF CANADA

This Society was incorporated in 1890 under the name of The Astronomical and Physical Society of Toronto, and assumed its present name in 1903.

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